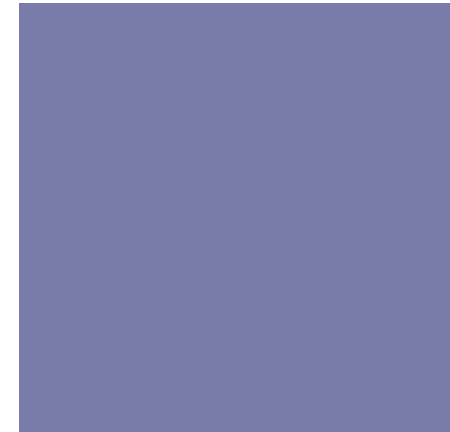
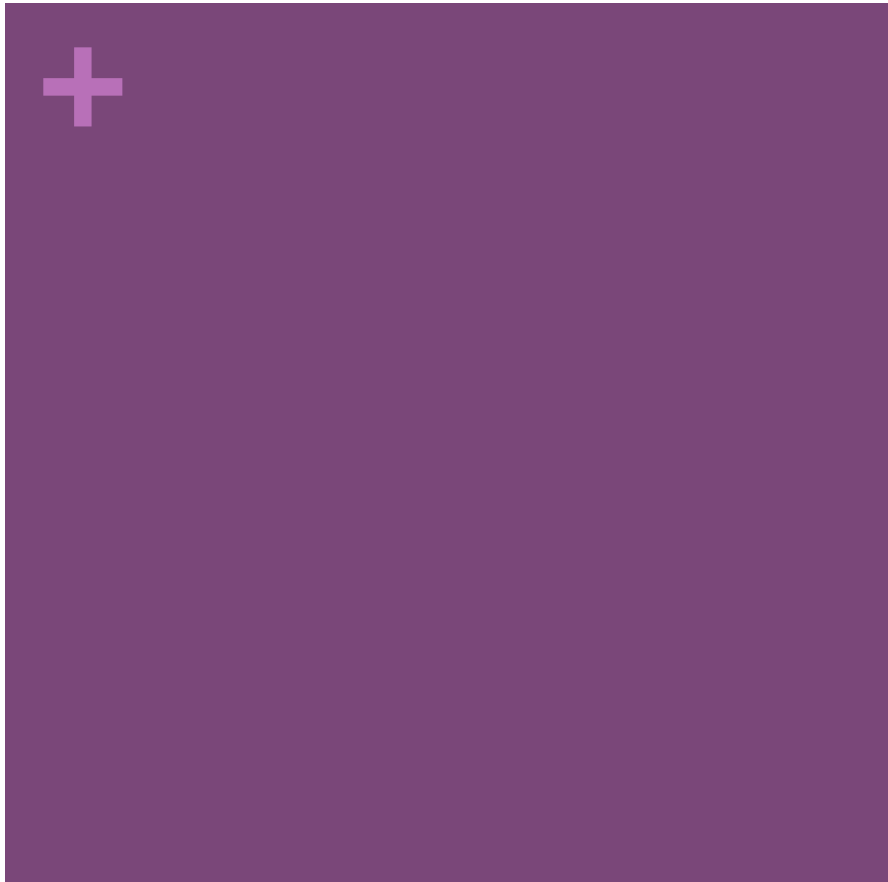


Experimental aspects: synchrotron radiation, beamlines, detectors, measurement modes geometry, sample preparation methods

Hiroyuki Oyanagi

Photonics Research Institute, AIST, Japan

IUCr 2011 XAFS Tutorial for crystallographers and beginners
Madrid, August 22, 2011

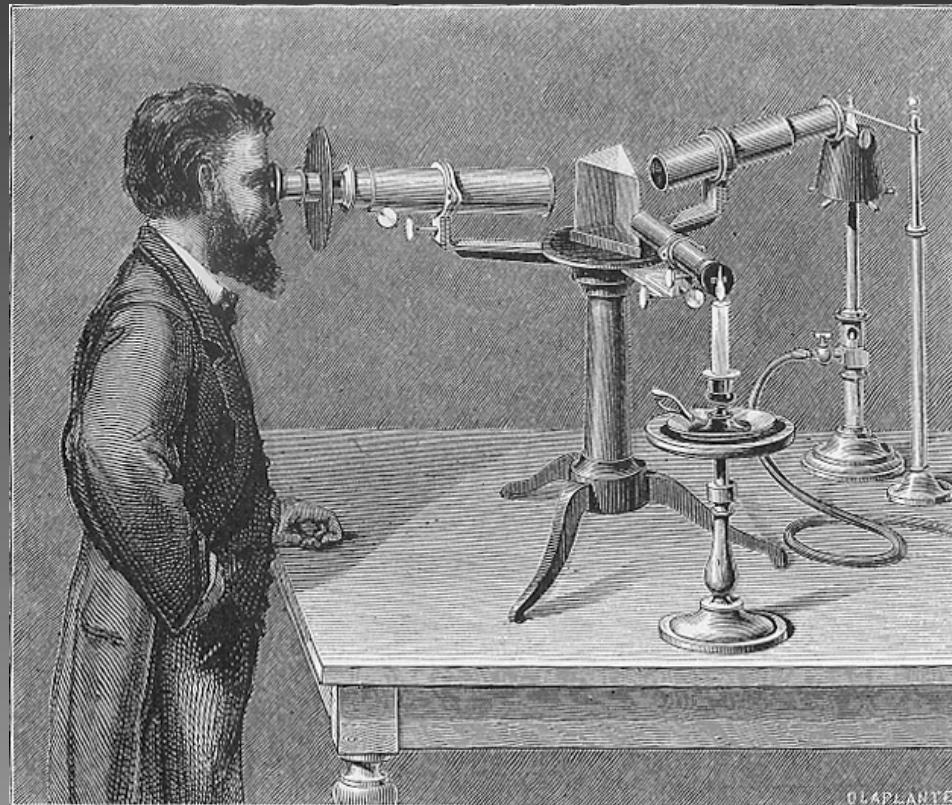


Introduction

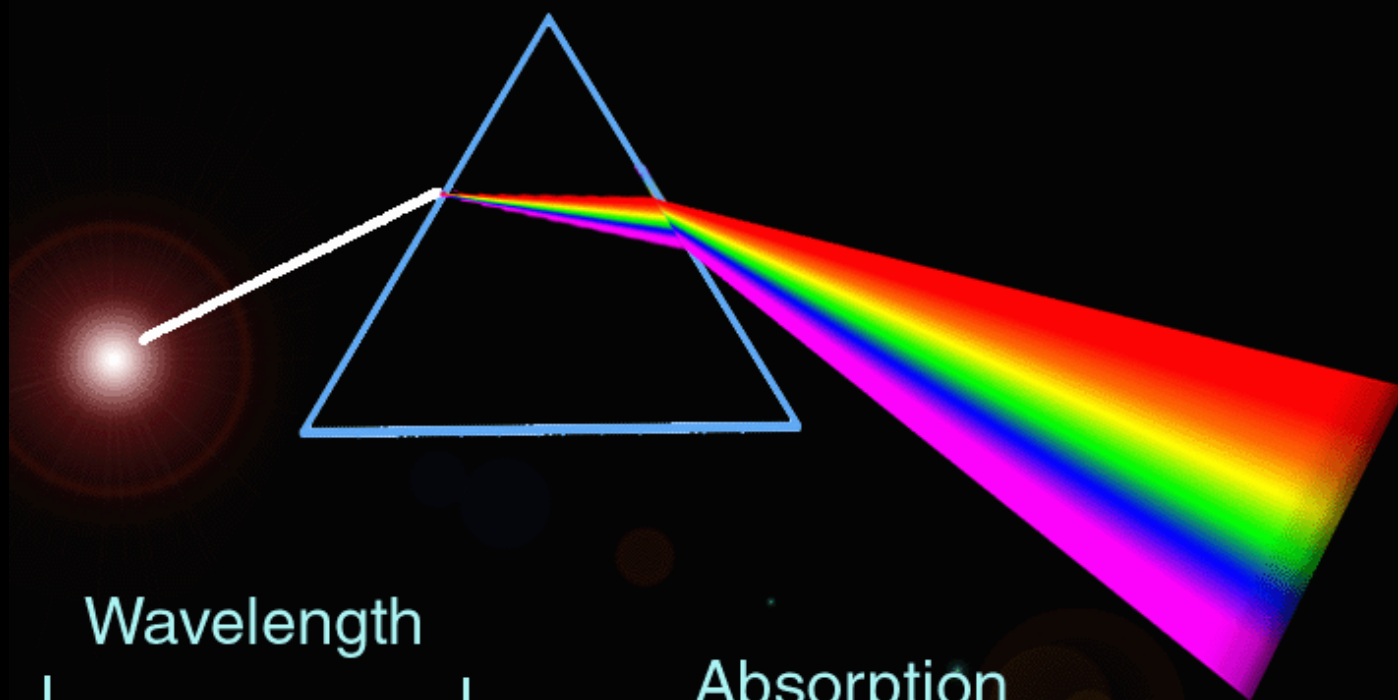
X-ray absorption spectroscopy
-brief description

What is spectroscopy?

The study of molecular structure and dynamics through the absorption, emission, and scattering of light



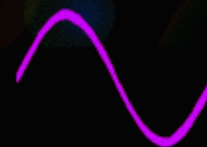
Astronomers used the “spectroscope”, to observe atomic spectra. Norman Lockyer found helium in the solar spectrum in 1868.



Wavelength



Absorption
lines

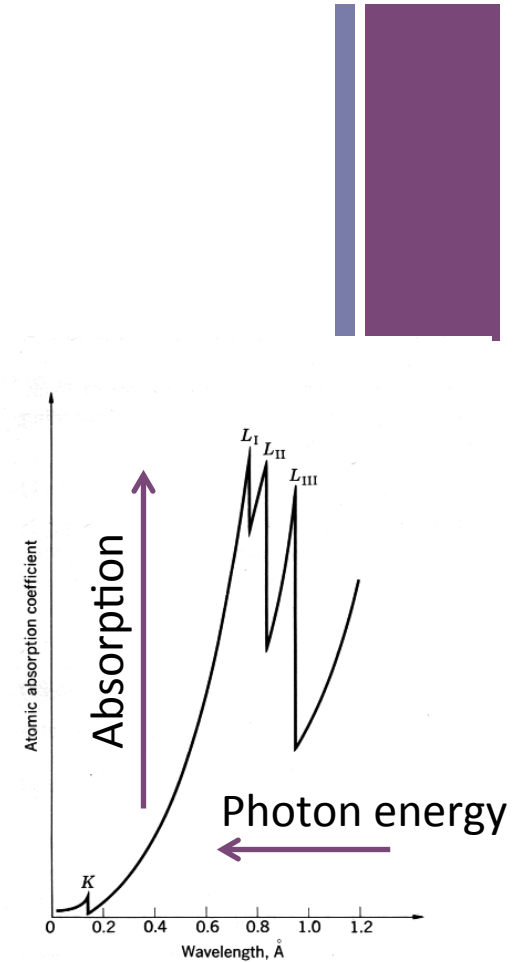
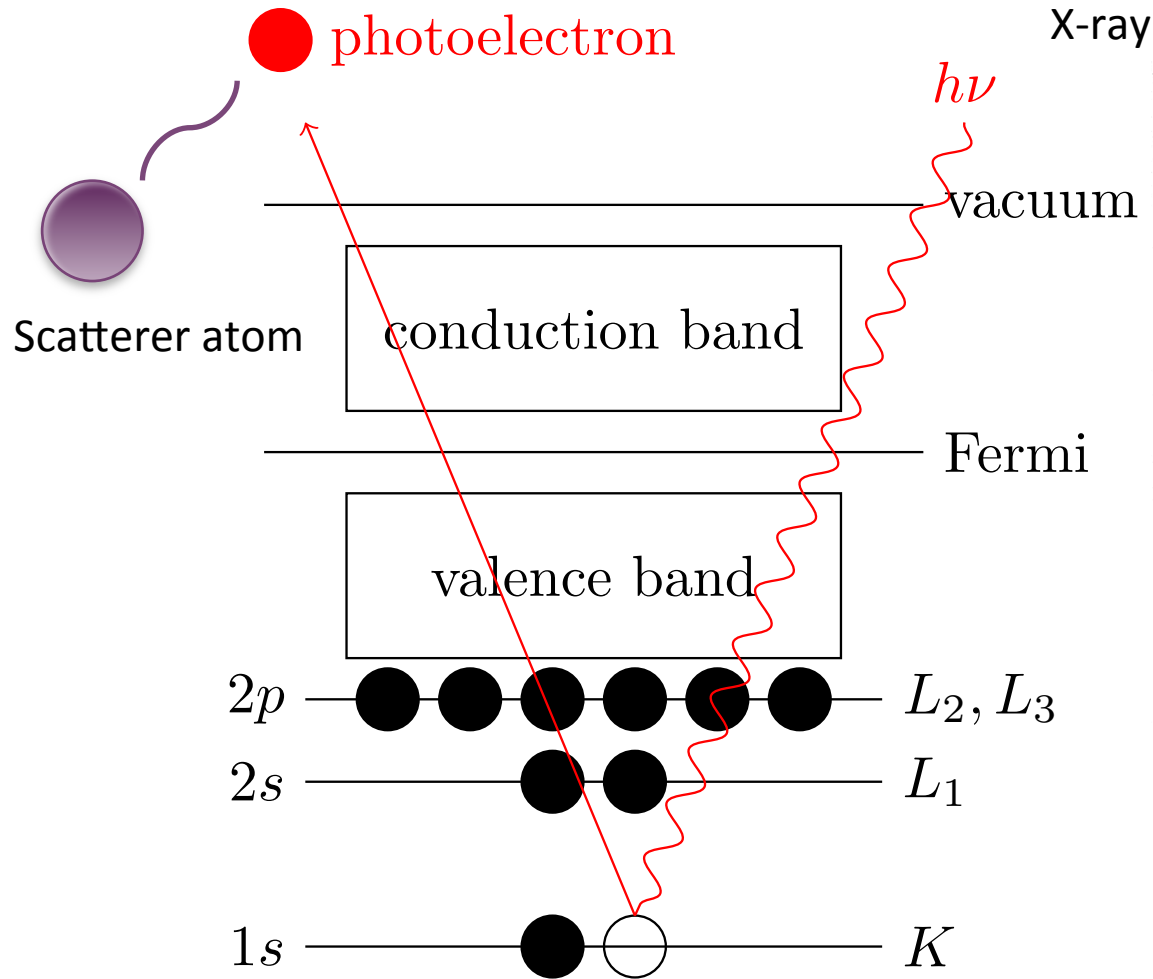


Hydrogen
H

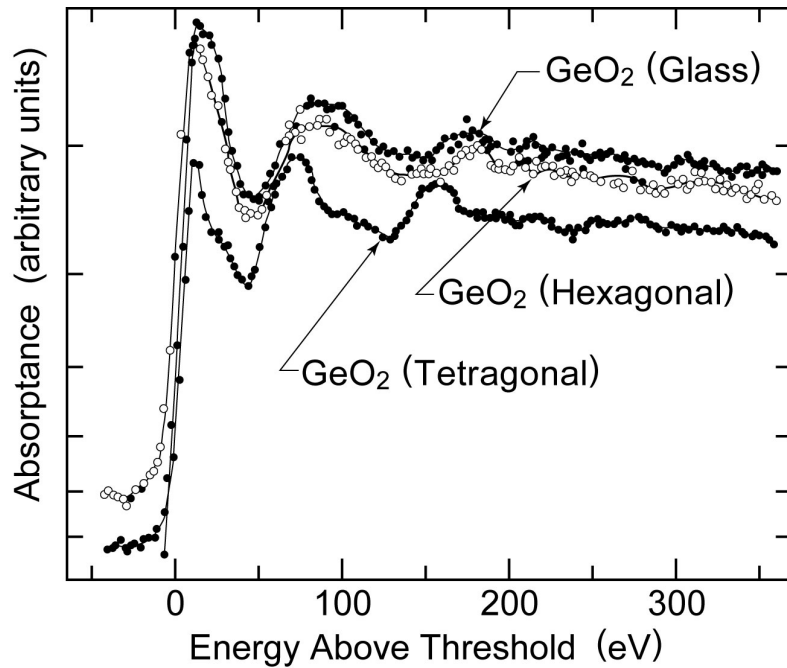
Sodium
Na

Calcium
Ca

+ X-ray absorption spectroscopy -schematic presentation

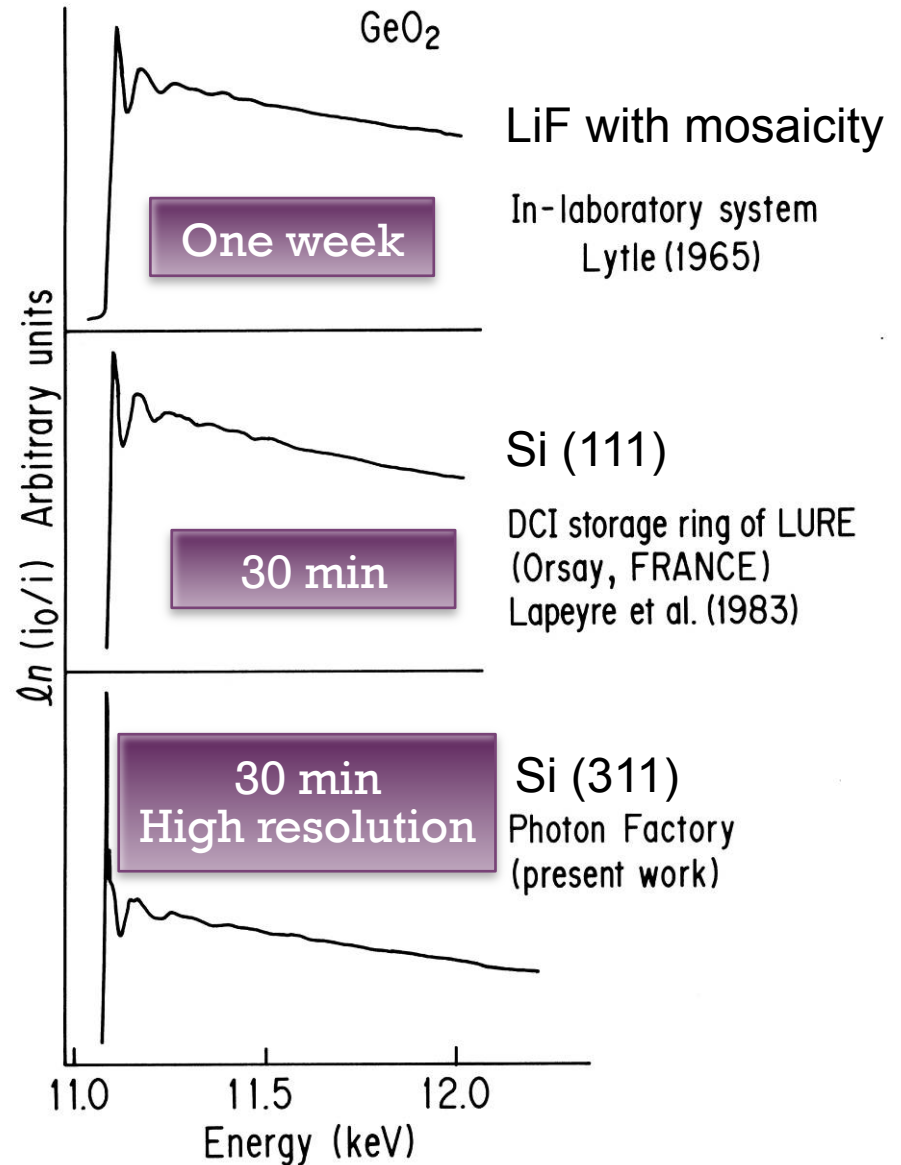


+ Power of Synchrotron radiation



Nelson et al., Phys. Rev. 127, 2025 (1962).
1 week for one spectrum, tube x-ray source + diffractometer

Synchrotron radiation changed quality & Quality of XAS



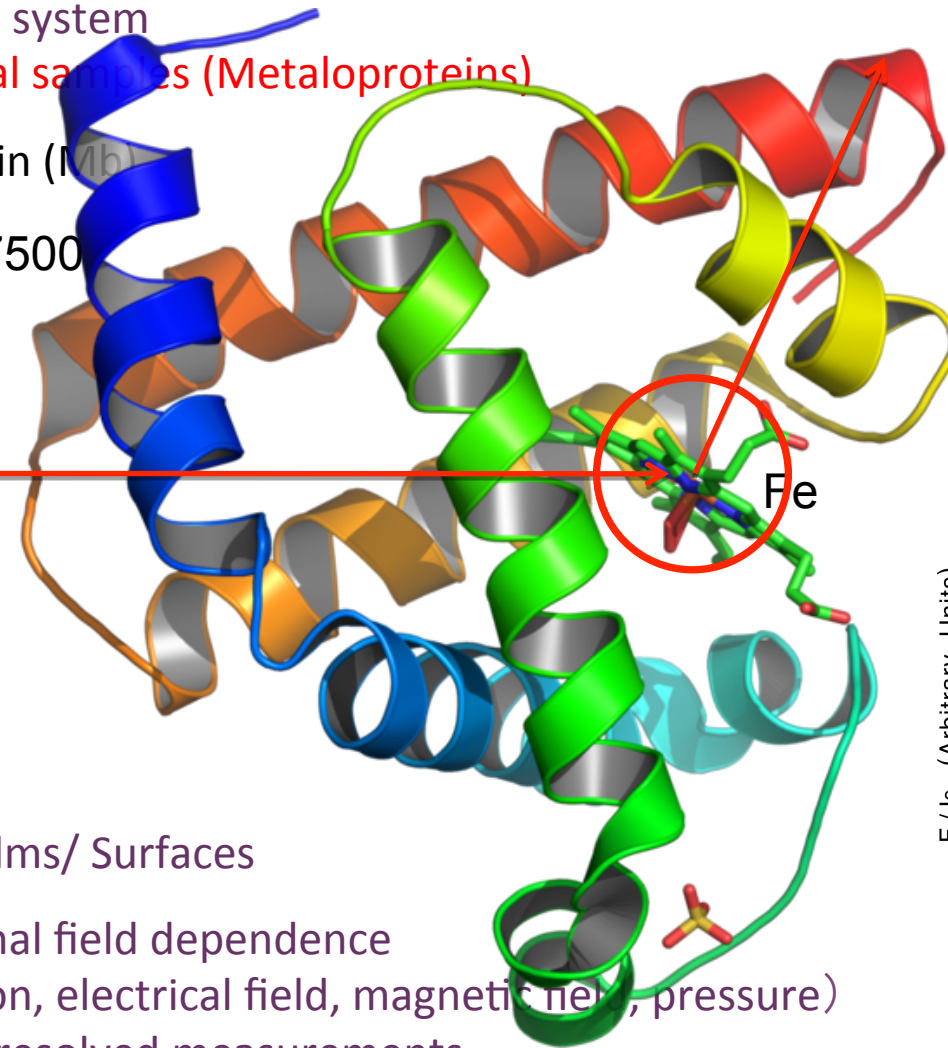
+ Fluorescence-XAS probing a dilute system

A. Dilute system
Biological samples (Metalloproteins)

Myoglobin (Mb)

1Fe/ 17500

X-ray



Fluorescence x-ray

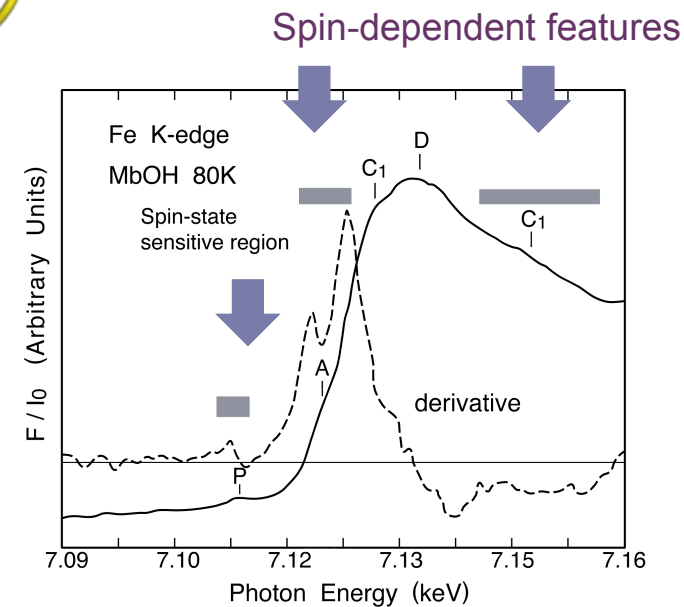
- **SPIN-STATES**
- **LOCAL STRUCTURE**
- **FUNCTION**

B. Thin films/ Surfaces

C. External field dependence
(photon, electrical field, magnetic field, pressure)

D. Time resolved measurements

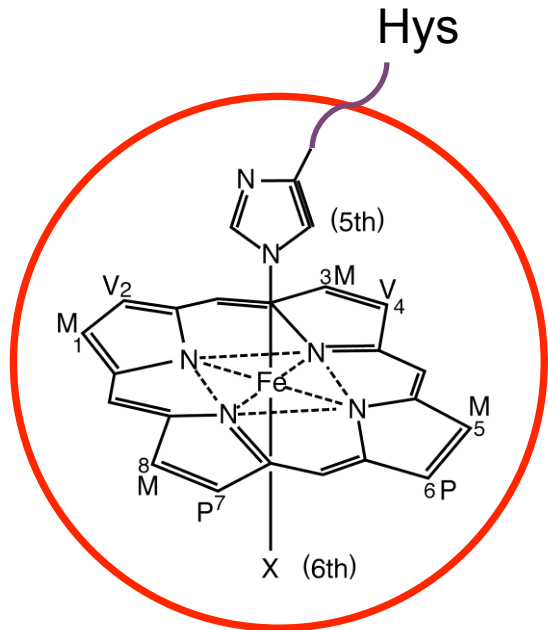
E. Small volume, nanocrystals, solutions



Spin-dependent XANES of MbOH

Oyanagi et al., J. Phys. Soc. Jpn. 56, 1987, 3381.

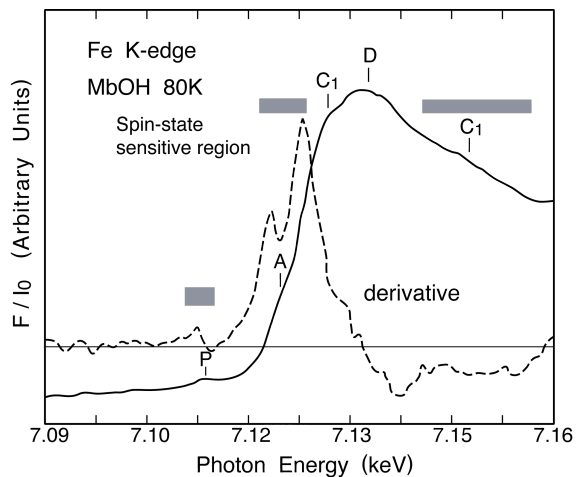
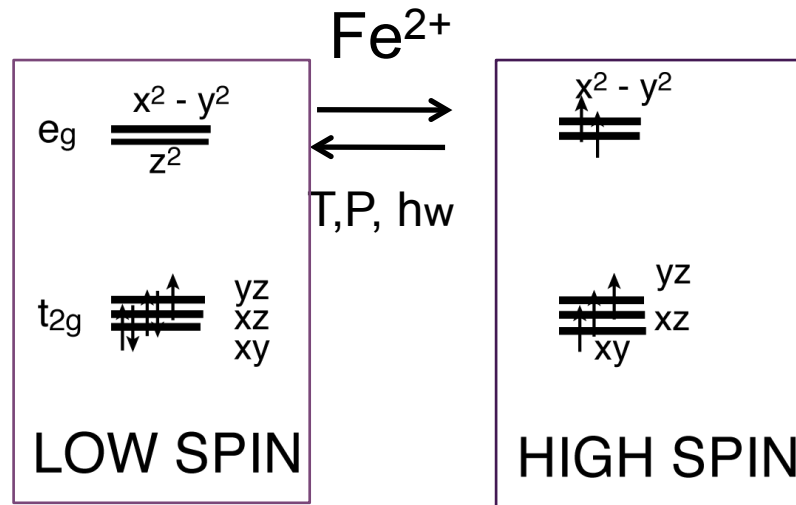
+ Local structure and spin states in MbOH



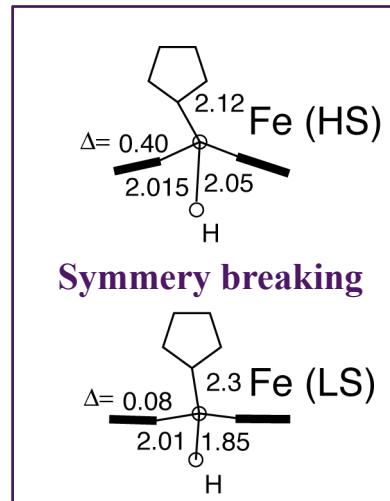
Fe's 5th bound to Hs, 6th is active site (oxygen bind)

$S = 1 / 2$

$S = 5 / 2$



Oyanagi et al., J. Phys. Soc. Jpn. 56, 1987, 3381.

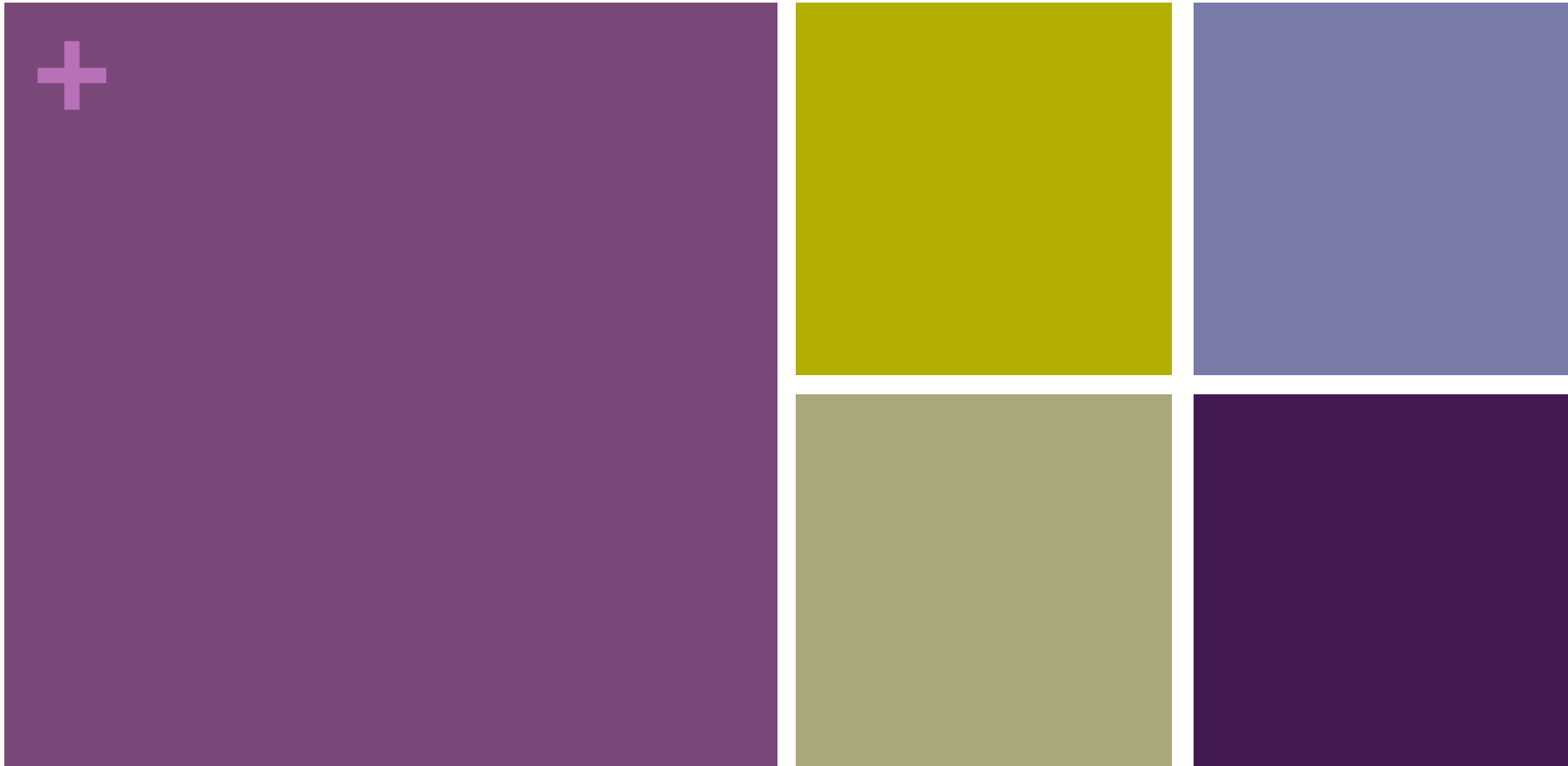


In HS, Fe is popped out by 0.4Å

In LS, Fe in the heme plane

Spin-dependent full multiple scattering
 S. Della Longa et al. J. Biol. Chem. 272, 21025 (1996).

Symmetry breaking to lower ligand field



Synchrotron radiation

Donuts proliferating the world
What are they?

+ Synchrotron radiation

3rd generation (3G) synchrotron radiation facilities

Started from a “mega” facility

Now proliferating as a “compact” machine



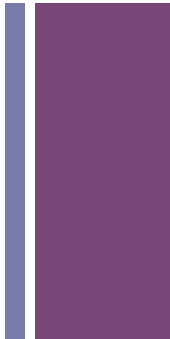
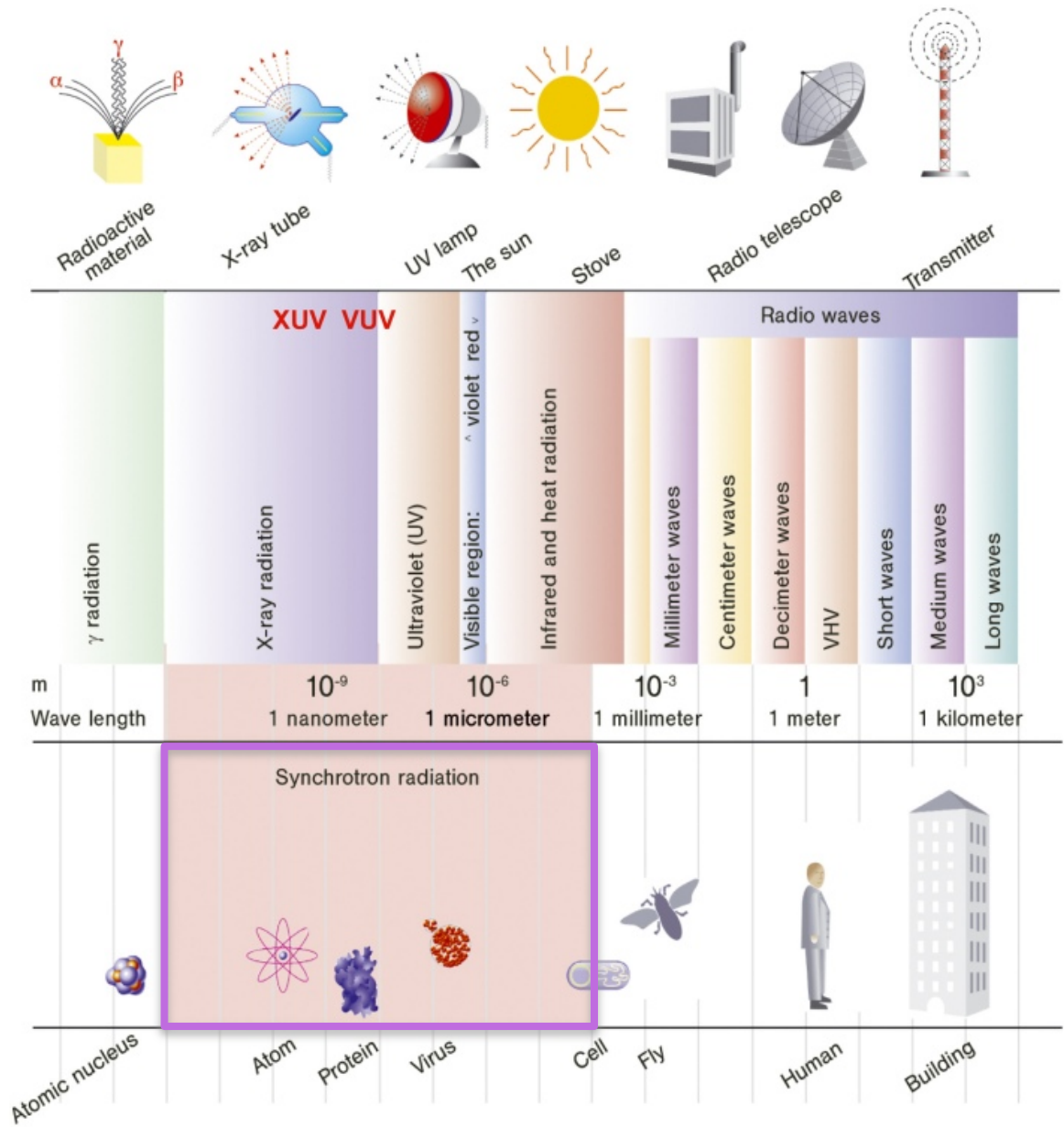
ESRF

ALBA



Apple mother ship

+ Wavelength and object size



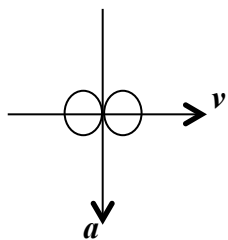
+ Synchrotron radiation –relativistic radiation

Normal radiation ($\beta \approx 0$) and relativistic radiation ($\beta \approx 1$)

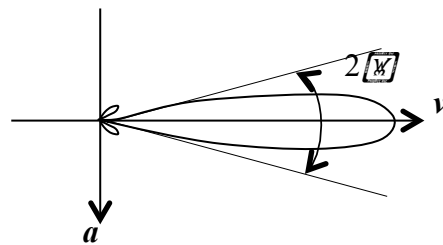
$$\beta \approx 0 \quad dP / d\Omega \approx 1 - \sin^2 \theta$$

$$\beta \approx 1 \quad dP / d\Omega \approx 1 / (1 - \cos \theta)^3$$

$$\Psi \approx 1 / \gamma = \sqrt{1 - \beta^2}$$

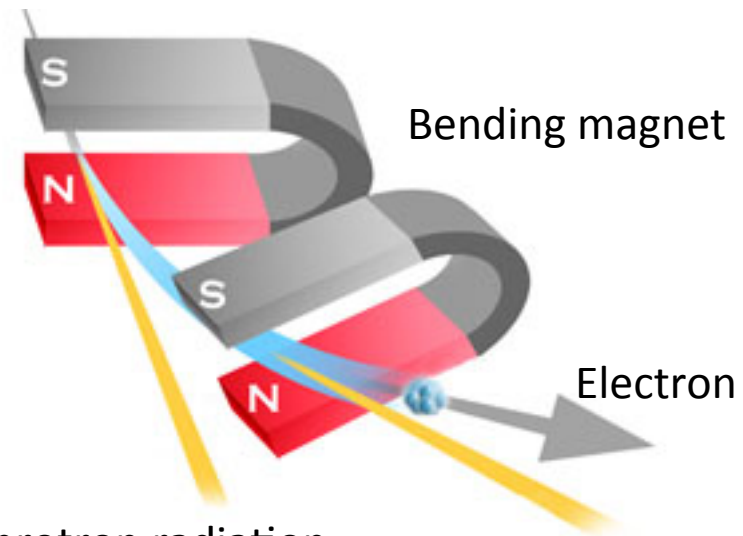


(a) $\beta \approx 0$



(b) $\beta \approx 1$

Bending magnet radiation



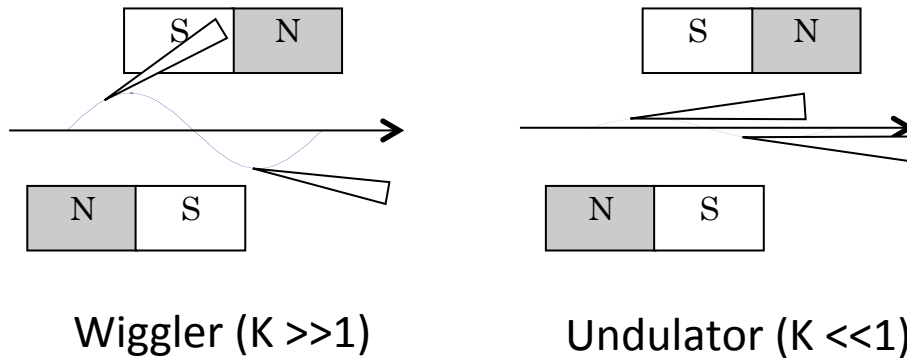
Synchrotron radiation

As electron velocity approaches c (that of light), radiation becomes highly directional, providing a bright white x-ray beam (synchrotron radiation)

+ Undulator and wiggler

$$K = \frac{eB_{\max}\lambda_o}{2\pi m_o c} = 0.934 B_{\max} [\text{T}] \lambda_o [\text{cm}]$$

B_{\max} Maximum flux density, λ_o Period length

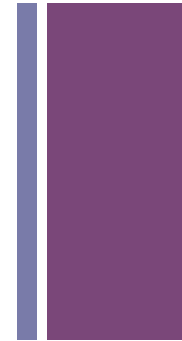


High magnetic field ($K \gg 1$)

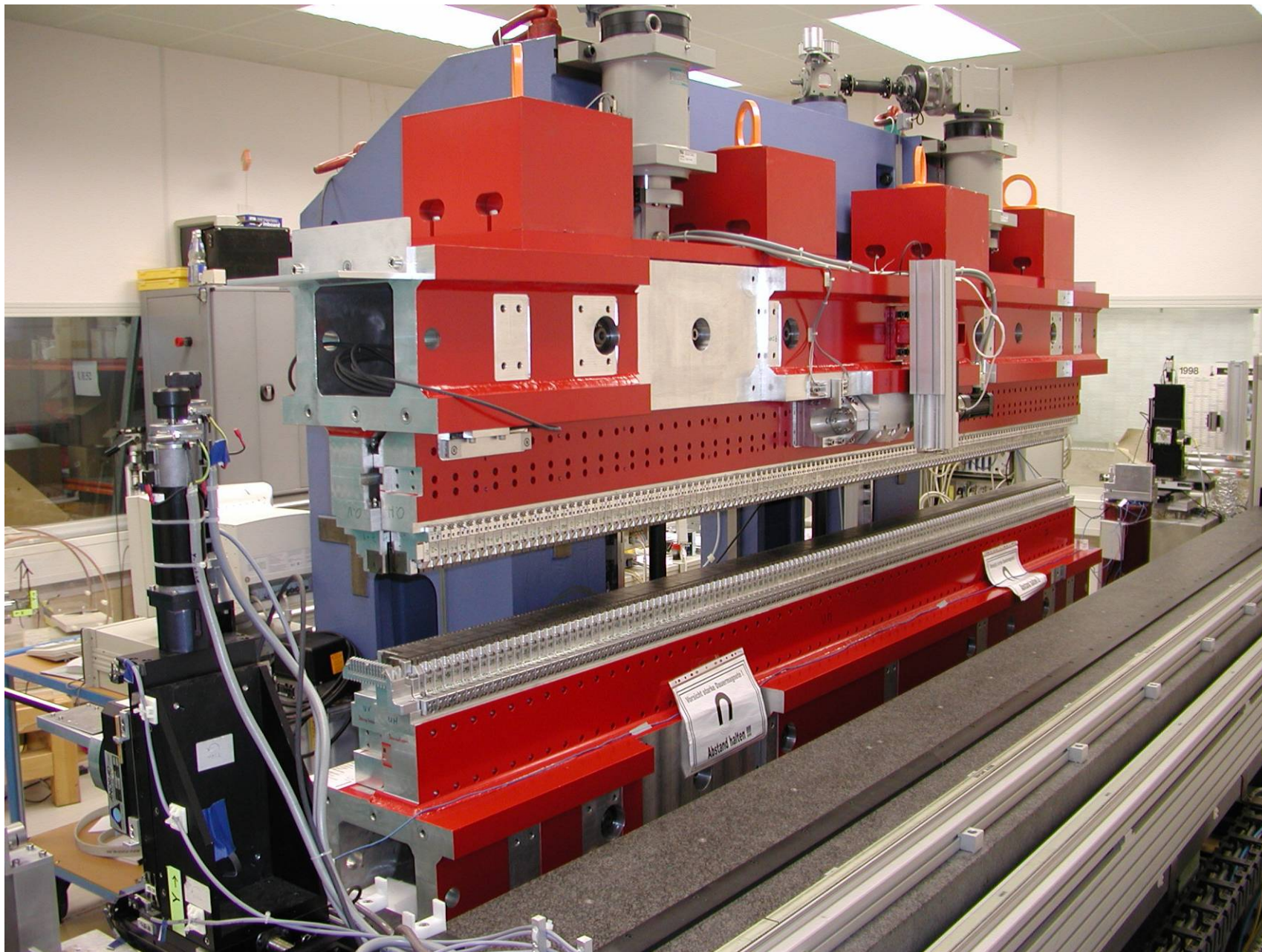
N-pole wiggler radiation enhances brilliance by N
Produces white x-ray but high heat load

Low magnetic field ($K \ll 1$)

Quasi-monochromatic high brilliance beam
Less heat load (high power density)



+ Undulator



+ Brilliance

Brilliance

10^{23}

Photons/s
/mrad²

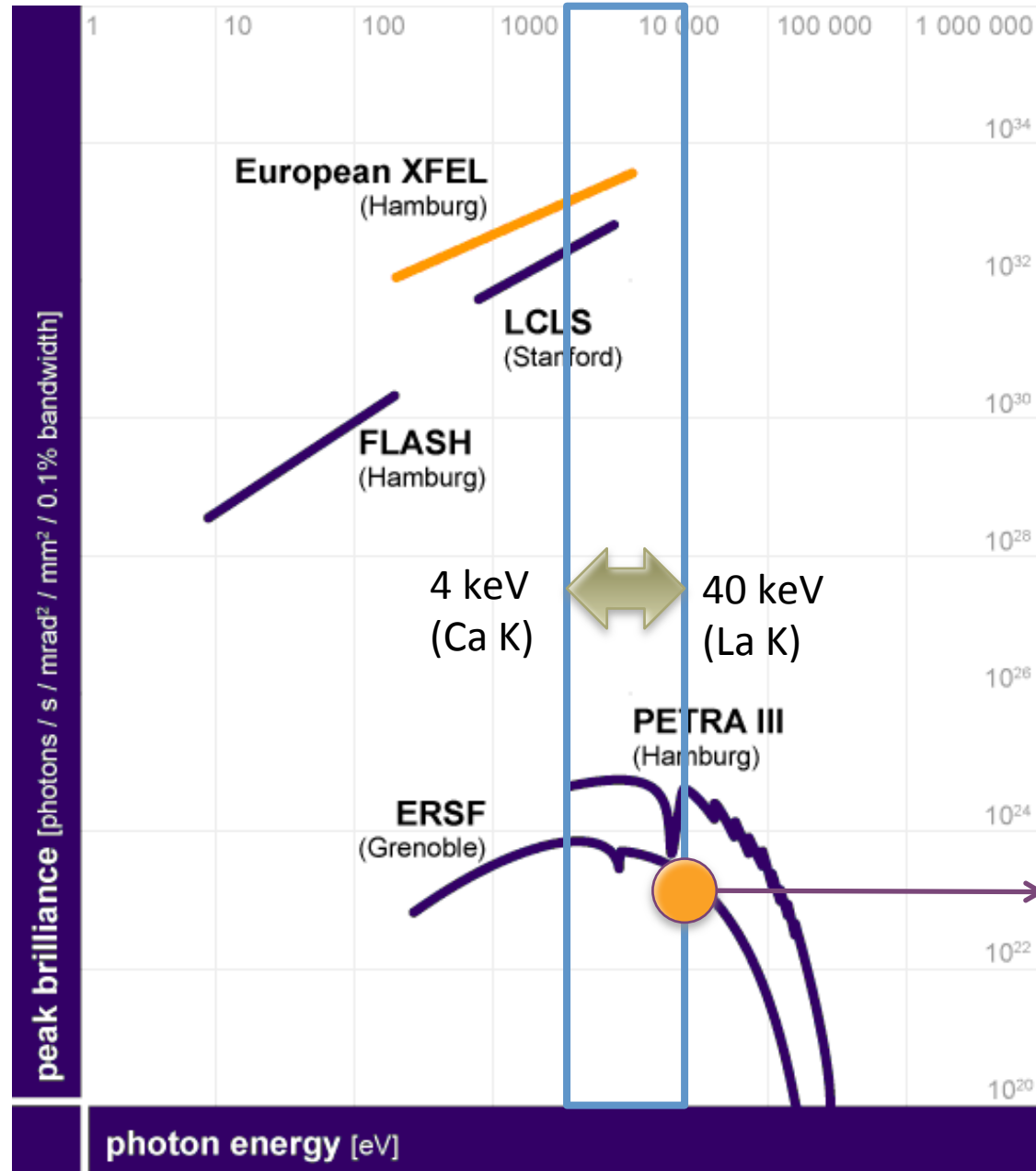
/mm²

/0.1% bandwidth

Flux

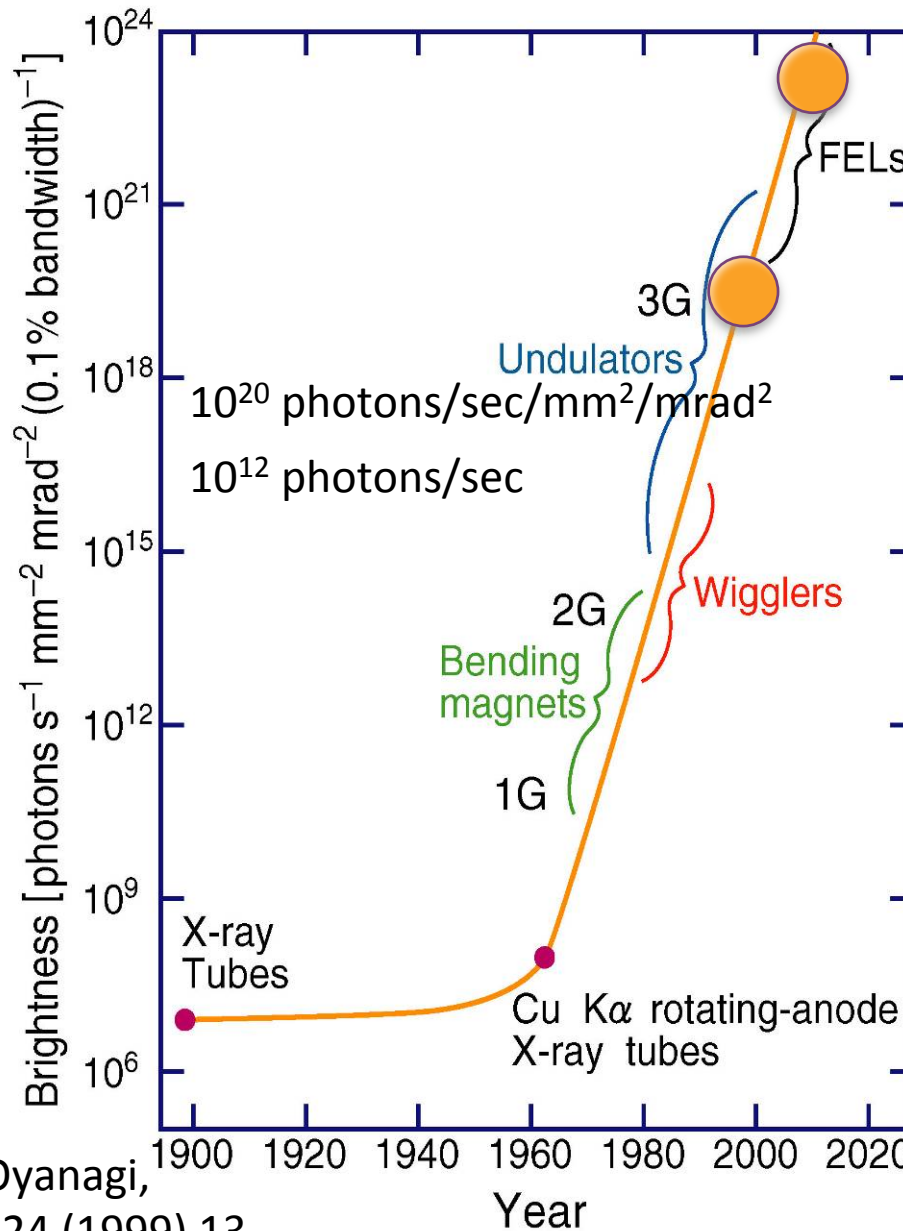
10^{12}

Photons/s



10^{23}

+ Moore's law in Synchrotron Radiation



We are here!

Note that exponential growth is due to successive inventions of different devices

10⁷ times brighter beam in 30 years

**NO
EXPONENTIAL IS
FOREVER...**

Gordon E. Moore

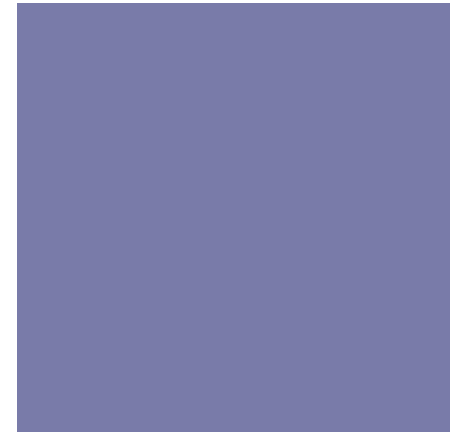
Montano and Oyanagi,
MRS BULLETIN 24 (1999) 13.



Omitted topics

VUV beamline

Soft x-ray beamline

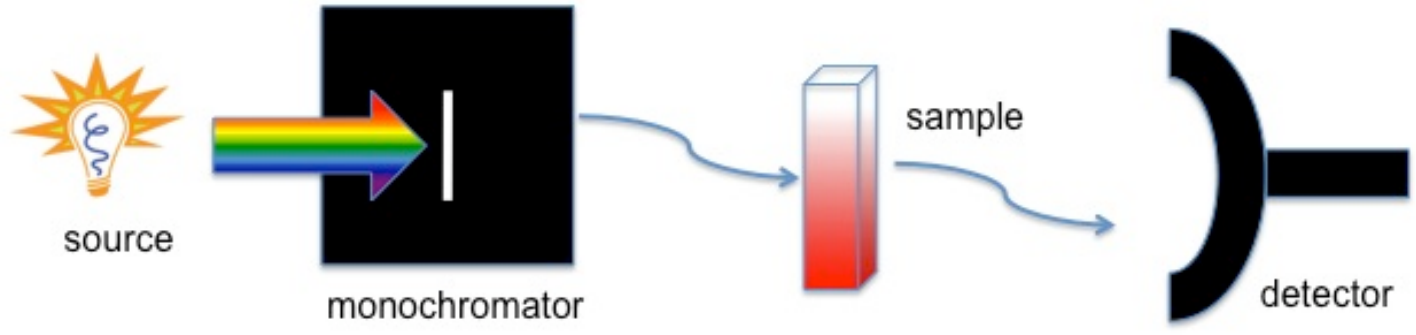


Beamline

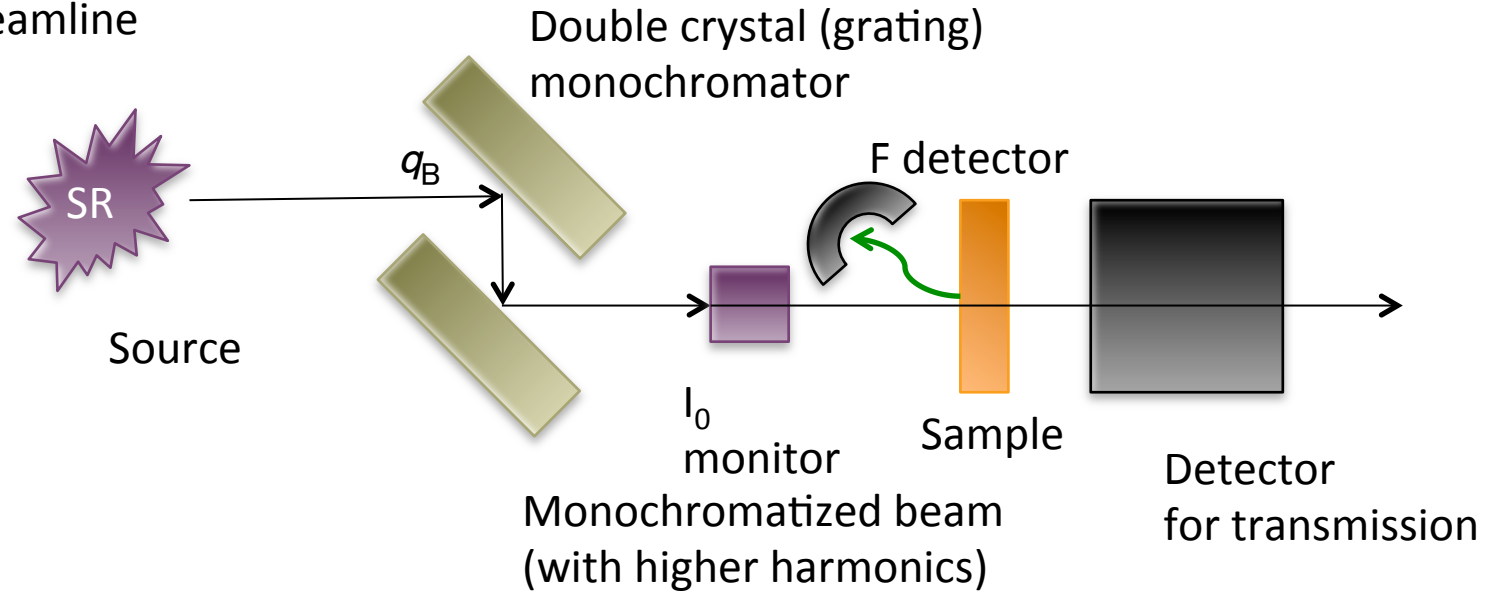
Monochromators and mirrors

+ XAS measurement –fundamental setup

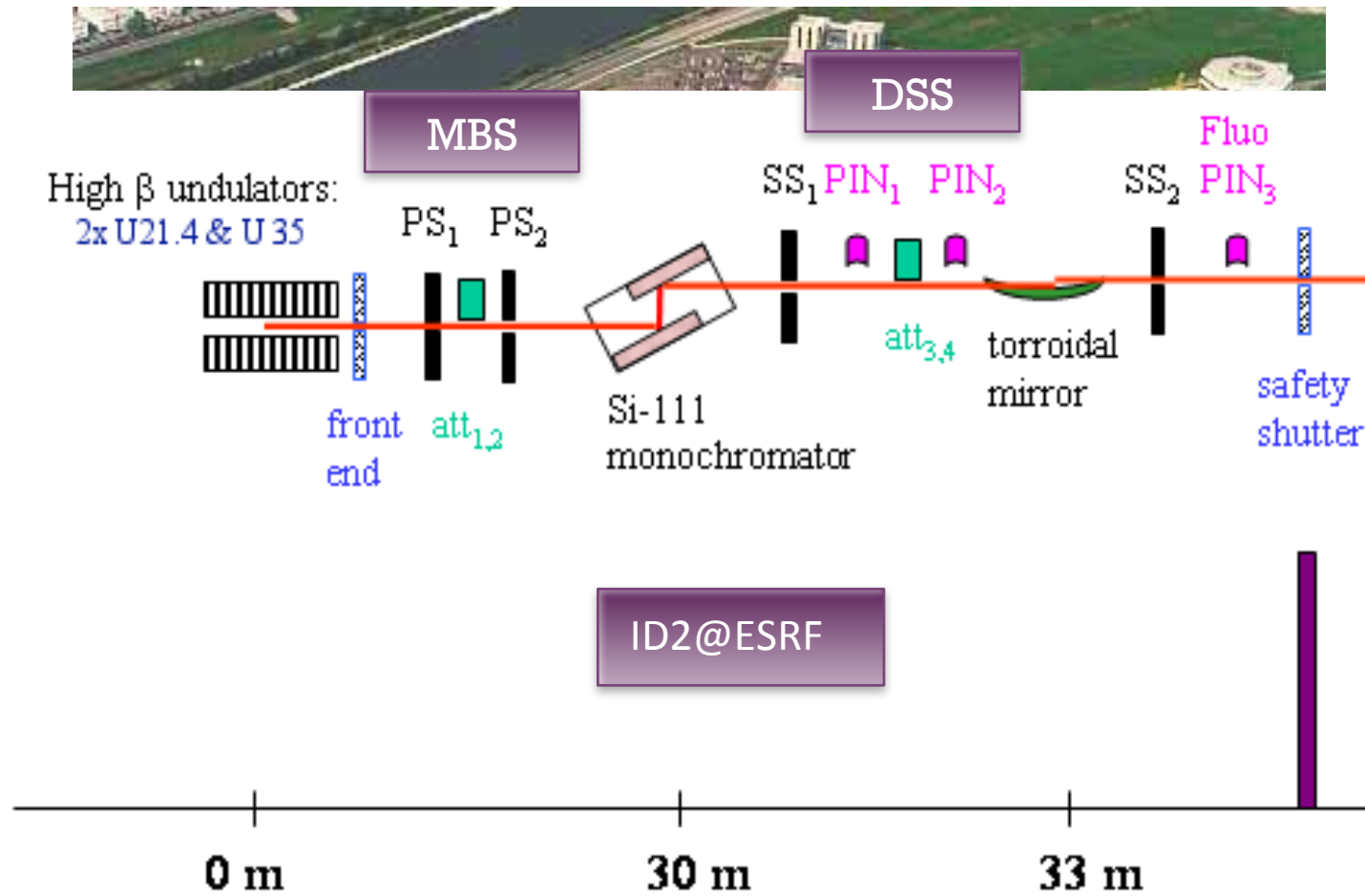
SINGLE BEAM



SR beamline



+ Storage ring and beam transport (beamline)

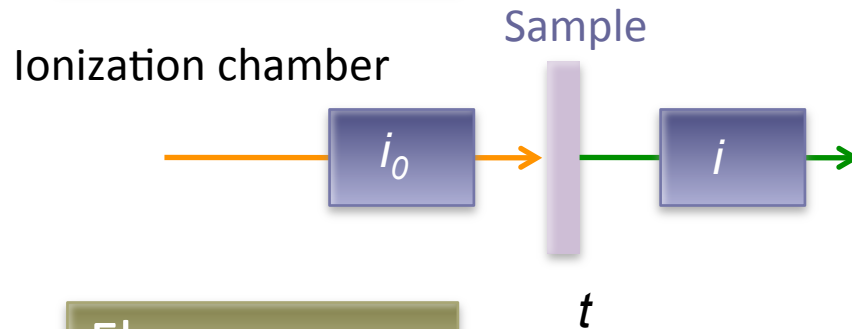


+ X-ray Absorption Spectroscopy -how to measure

XANES, EXAFS, ...

Most fundamental technique is a transmission mode

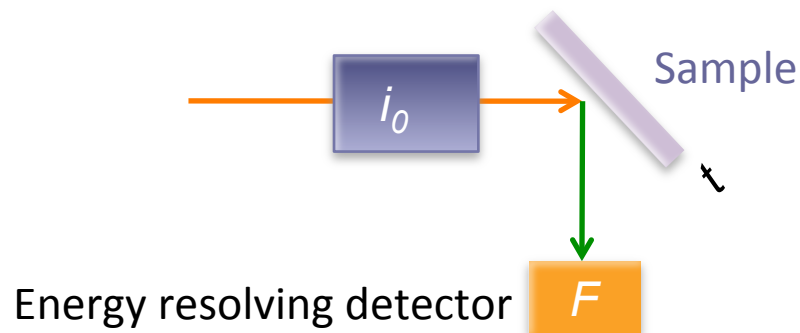
Transmission



$$mt(E) = \ln(i_0/i)$$

You measure attenuated beam intensity, that "exponentially" decreases

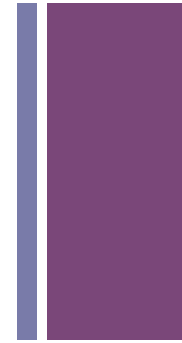
Fluorescence



$$mt(E) = F / i_0$$

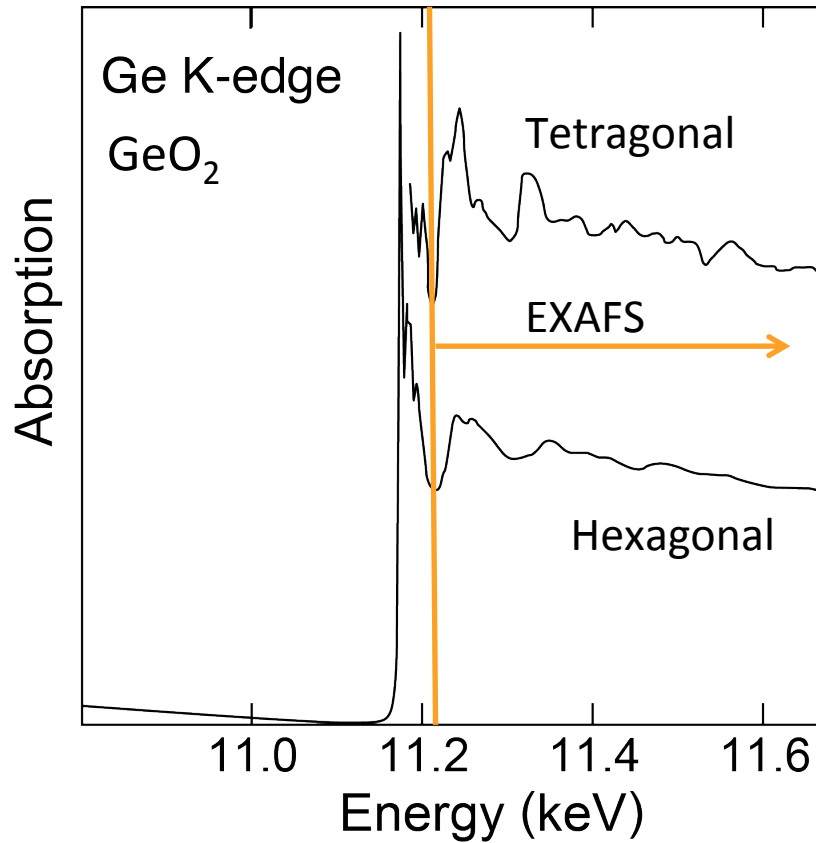
You measure emitted beam intensity
Which "linearly" proportional to conc.

Dilute system
Impurities, surfaces, thin films

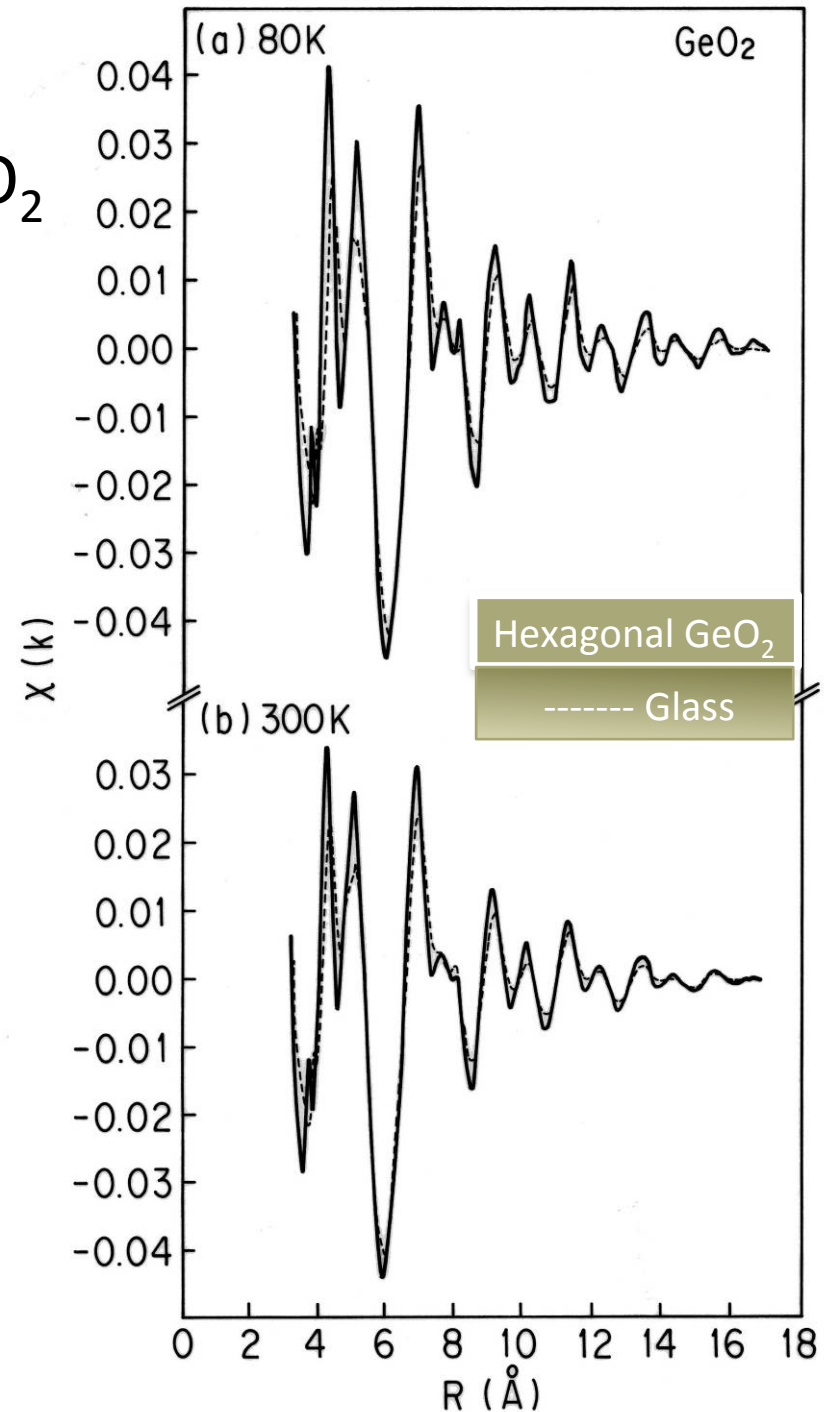


+ Example
- Crystalline and glassy GeO₂

Okuno et al.

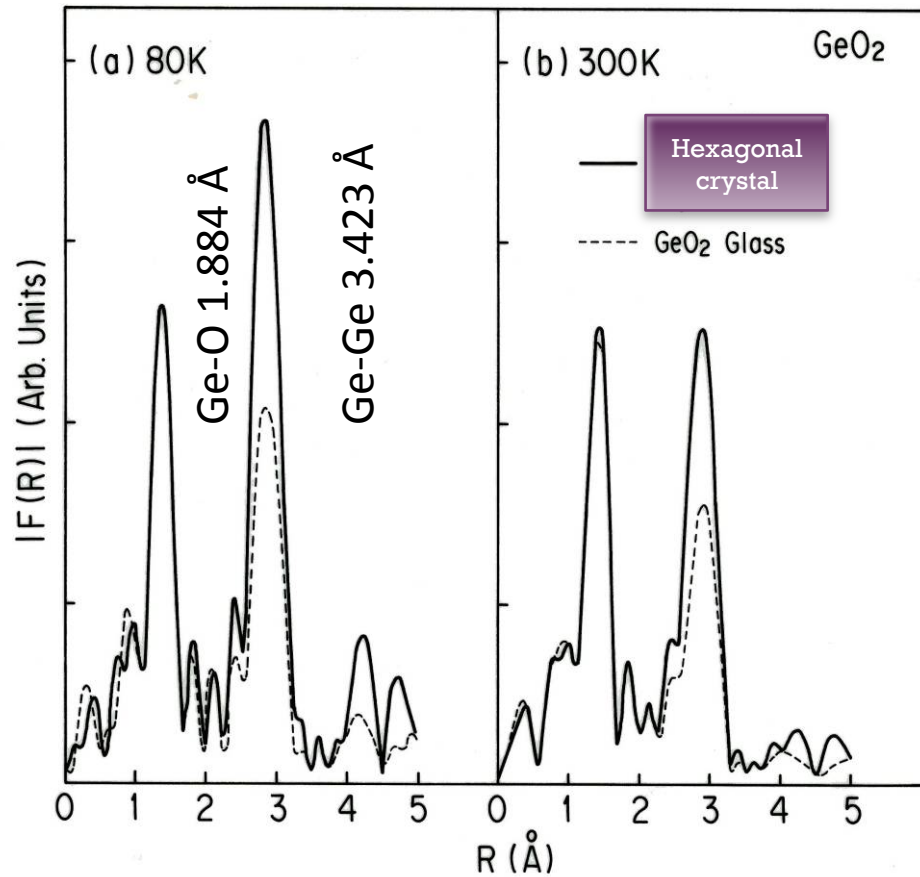
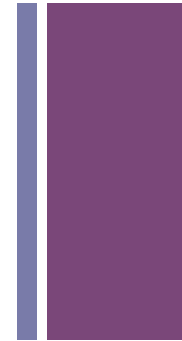


Si (311) monochromator, 10B@PF

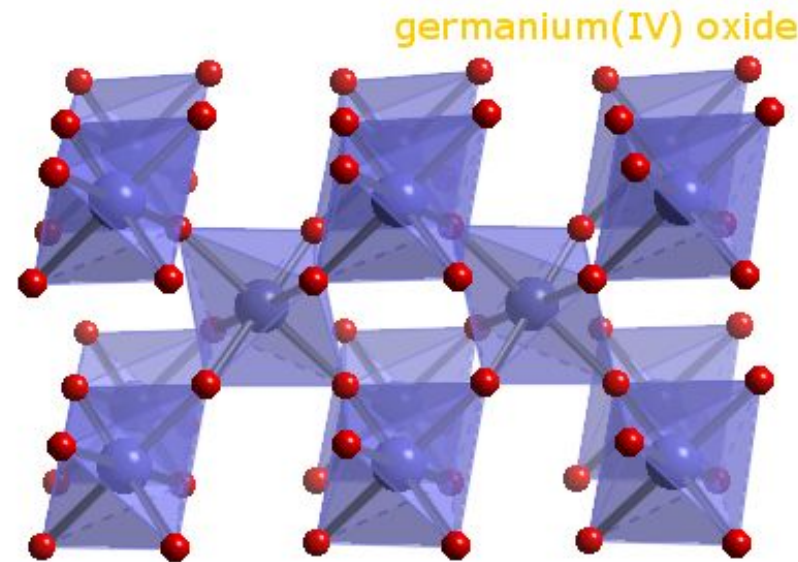


+ Fourier Transform -example

FT magnitude function for crystalline and glass GeO₂ Okuno et al.



Crystal structure (hexagonal)



Acta Cryst. 17, 842 (1964)

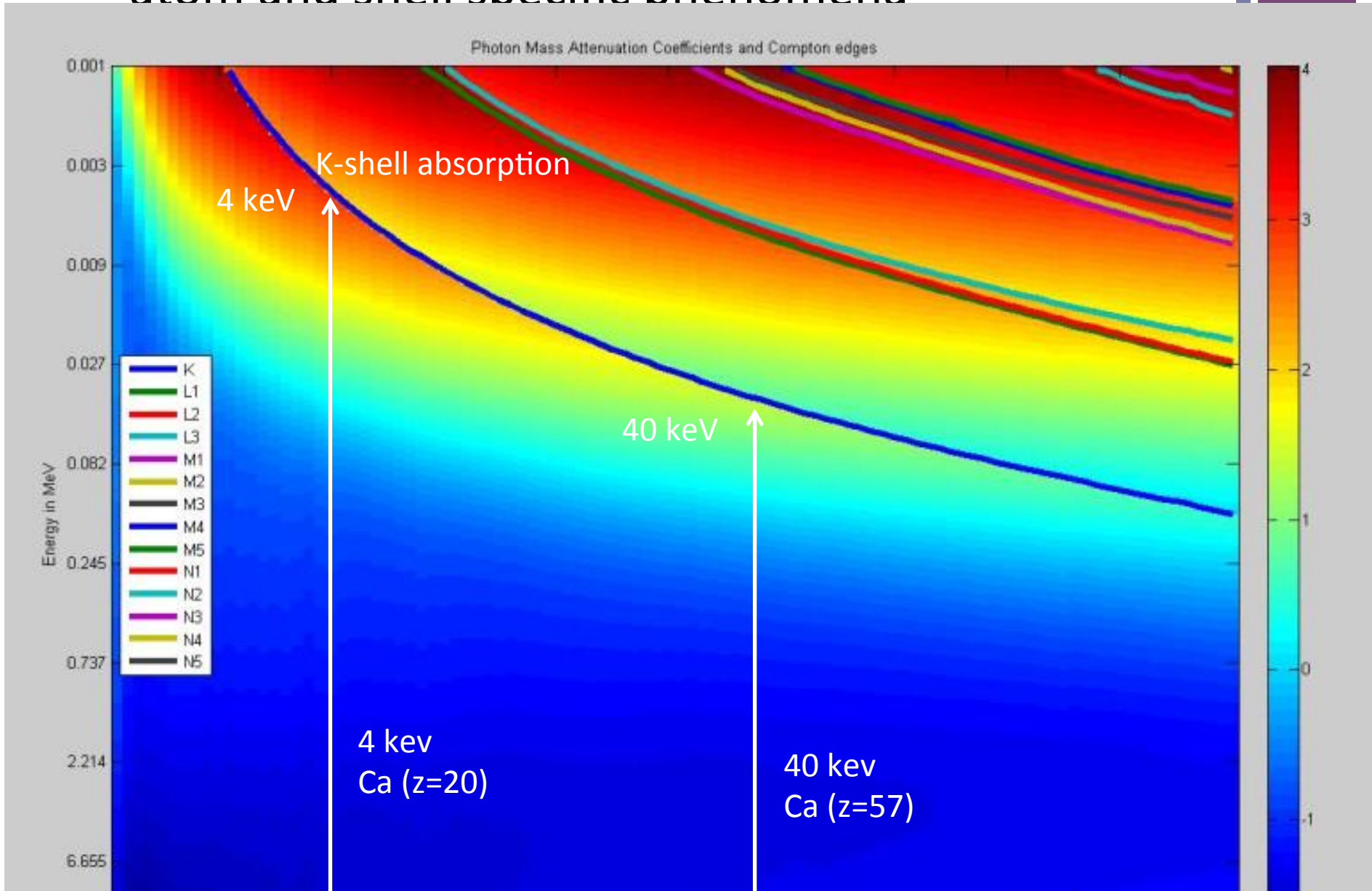
Acta Cryst. B27, 2133 (1971)

Glass structure

Short range order is close to the hexagonal crystal
Disorder in arrangement of GeO₄ units (connectivity)

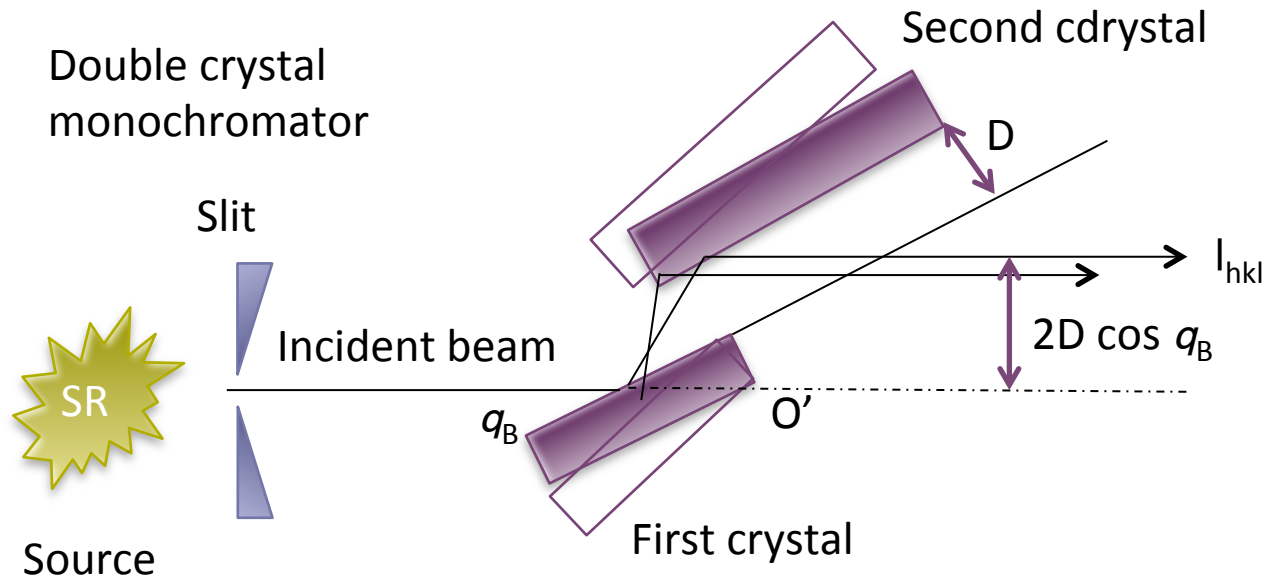
+ X-ray absorption

-atom and shell specific phenomena



+ Monochromator

Double crystal
monochromator



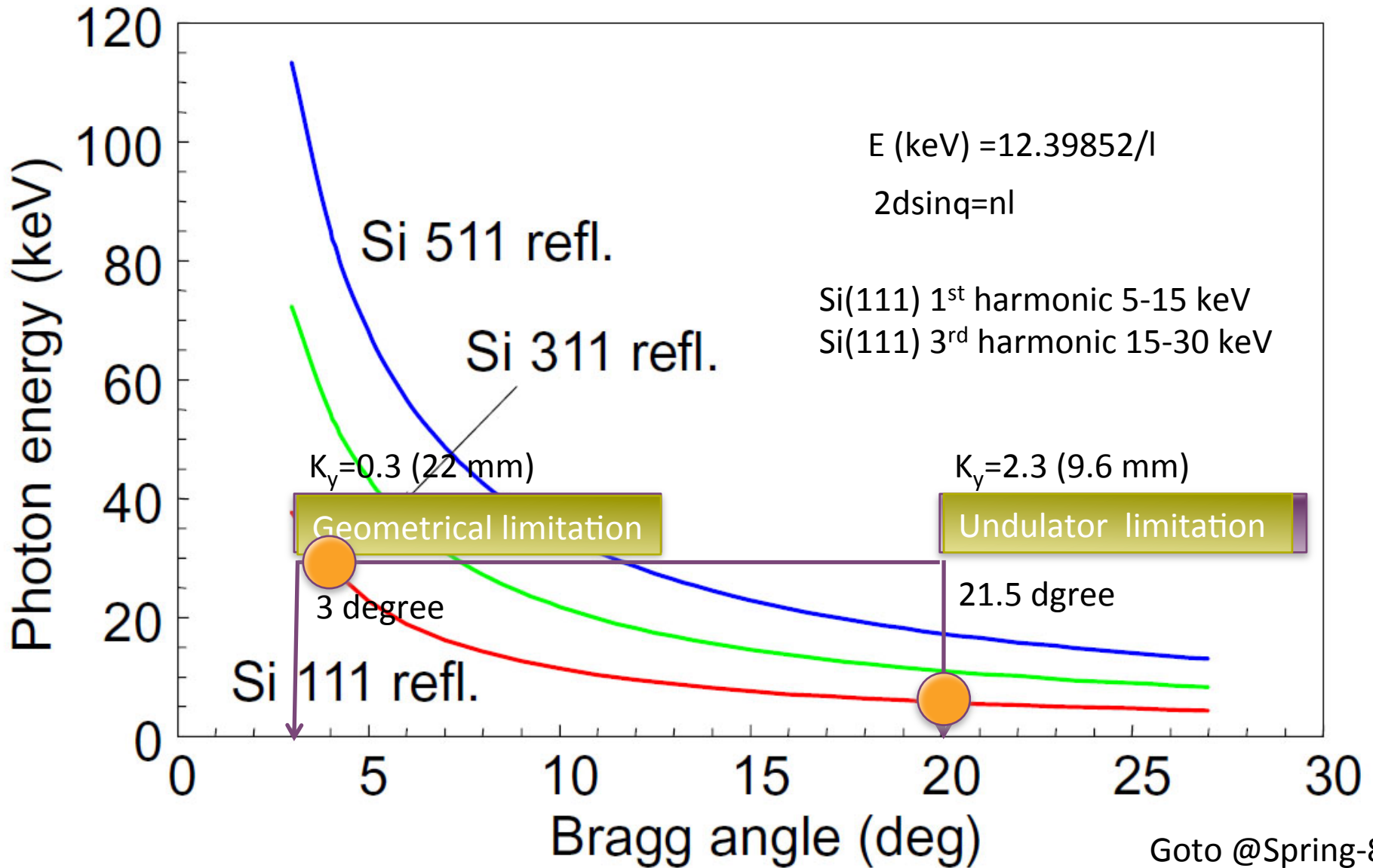
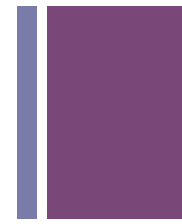
Energy scale

$$2d \sin q = n\lambda, \quad d = 3.13551 \text{ for Si(111)}$$

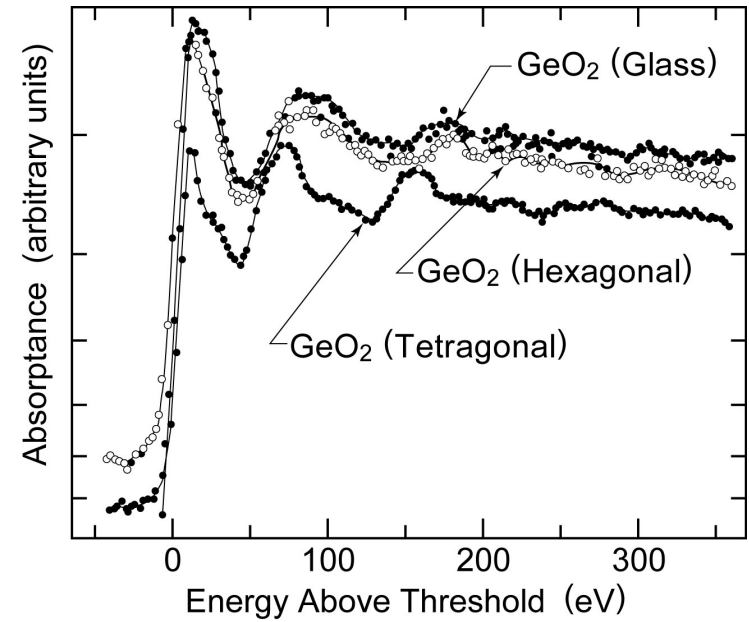
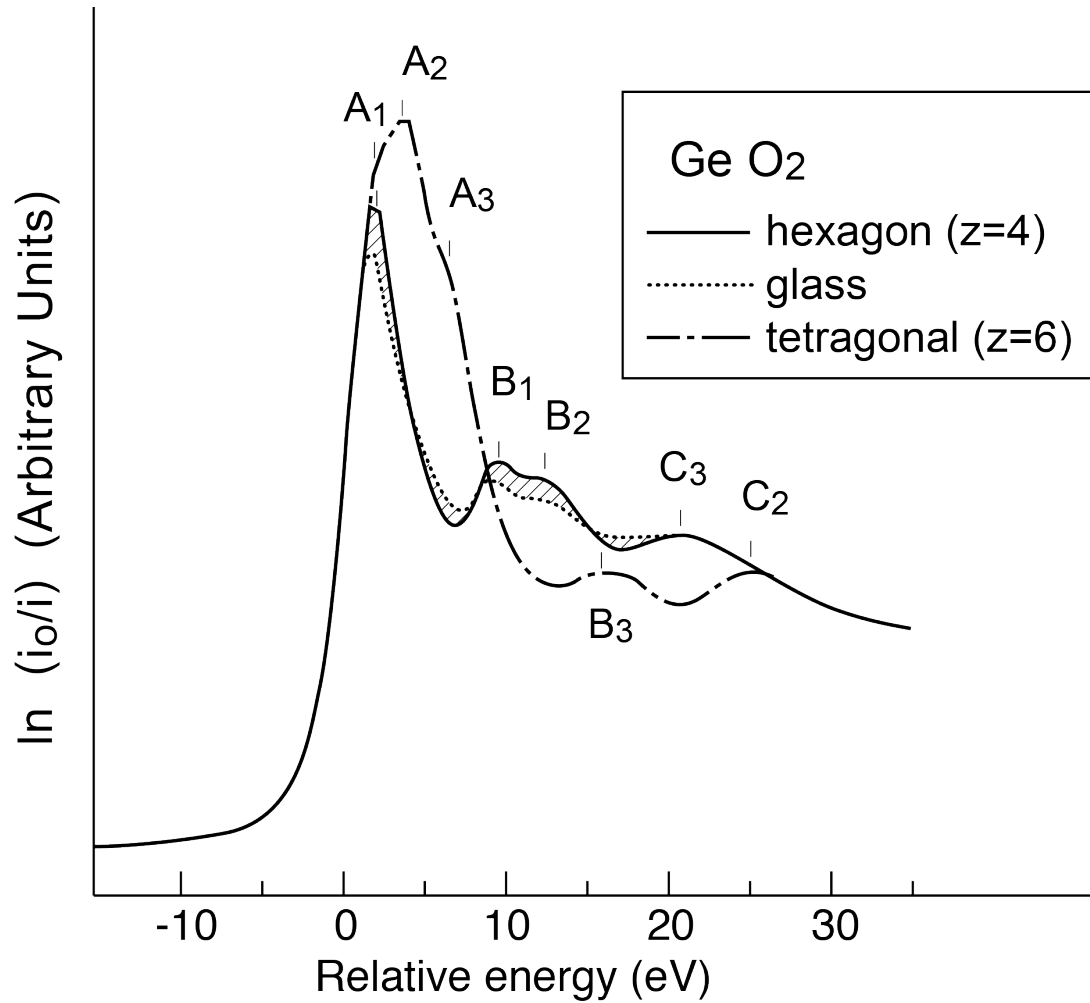
$$E \text{ (keV)} = 12.39852 / \lambda = 1.9771 / \sin q$$

Spacing (d -value)	Crystal plane
1.3578 Å	Si (400)
3.1356 Å	Si(111)
1.6376 Å	Si(311)
1.0452 Å	Si(511)

+ Monochromator –energy range



+ Example of XANES – why resolution?



Si (311) monochromator Okuno et al.

Nelson et al., Phys. Rev. 127, 2025 (1962).
1 week for one spectrum, tube x-ray source + diffractometer

Short range in glass sample is close to that of hexagonal crystal

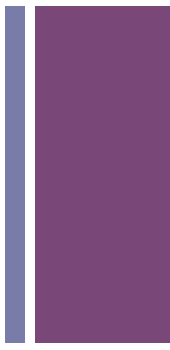
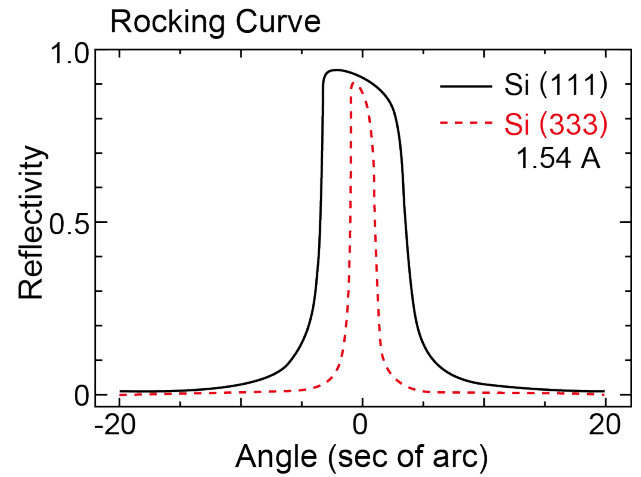
+ Energy resolution

Energy resolution $DE/E = \sqrt{(dq_g)^2 + (dq_w)^2} \cot \alpha_B$

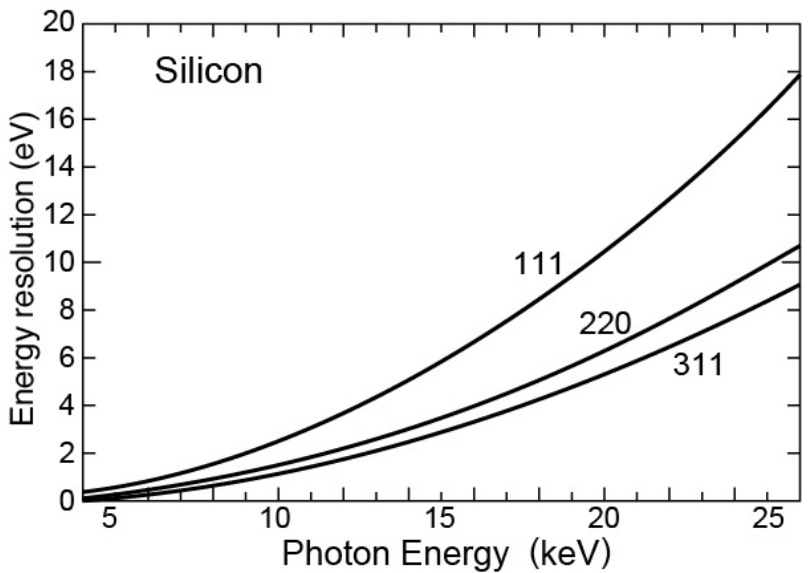
Geometrical resolution: dq_g

Darwin width: dq_w

$$dq_w = \frac{2e^3 |F|}{\rho m c^2 V \sin 2\alpha_B} = 8 \text{ sec. } (4 \times 10^{-5} \text{ rad}) \text{ (Si(111), 9keV)}$$



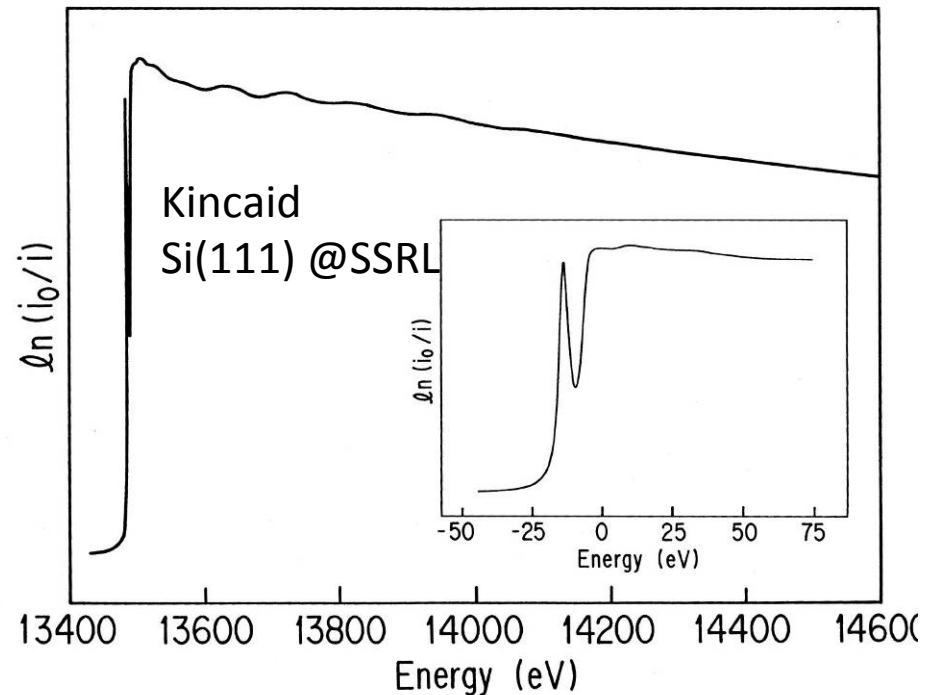
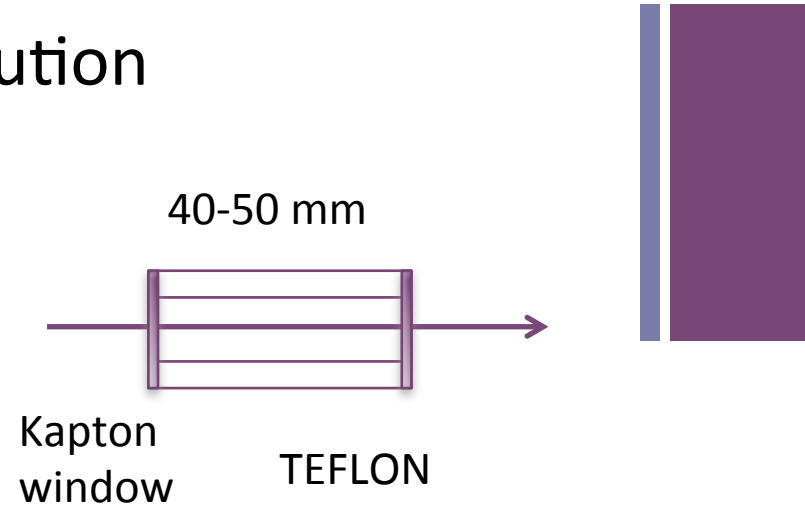
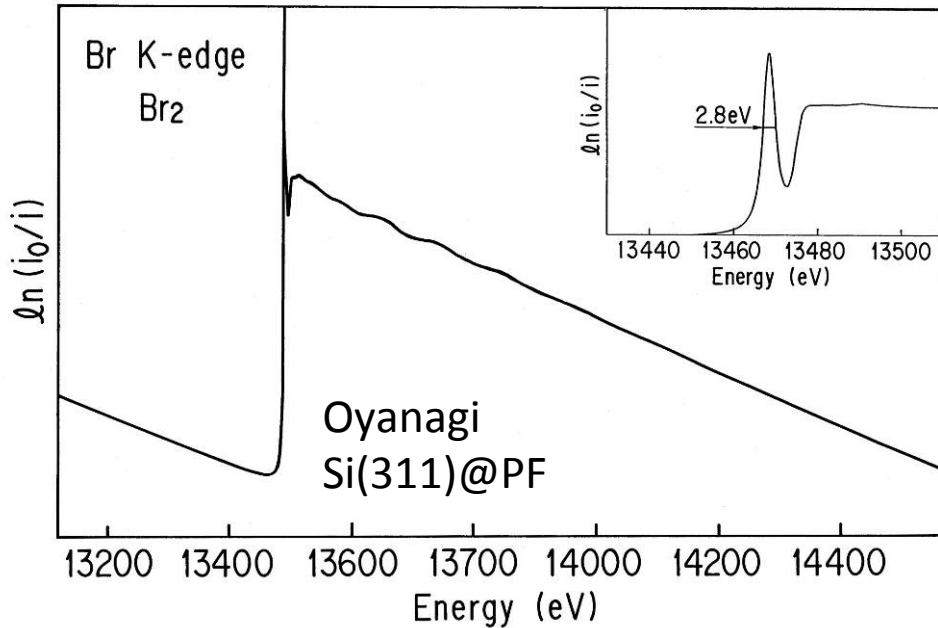
Energy resolution is given as a convolution of geometrical resolution and Darwin width



Guidelines

- a. Bragg-angle-dependent energy resolution degrades with the increase of energy
- b. Darwin width is smaller for high index planes, i.e. better resolution

+ Evidence for high energy resolution -bromine gas K-edge



How to evaluate the energy resolution

Br gas By K-edge (13.5 keV)

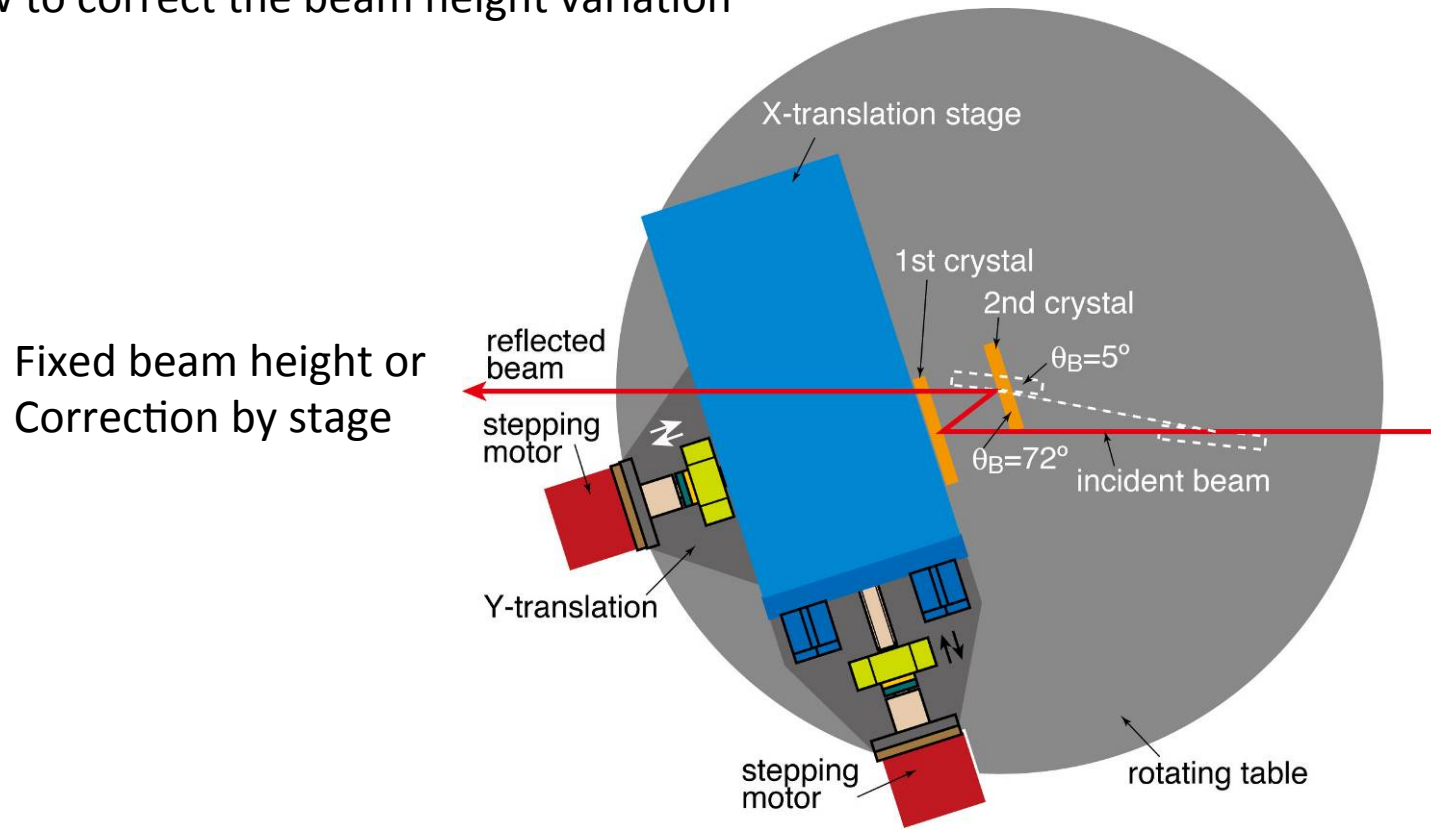
★Darwin width and angle-dependent term

★Abberation caused by a focusing mirror placed in front of monochromator degrades energy resolution T. Matsushita@SSRL

+ Monochromator -mechanism

Rotating a double crystal results in beam height change

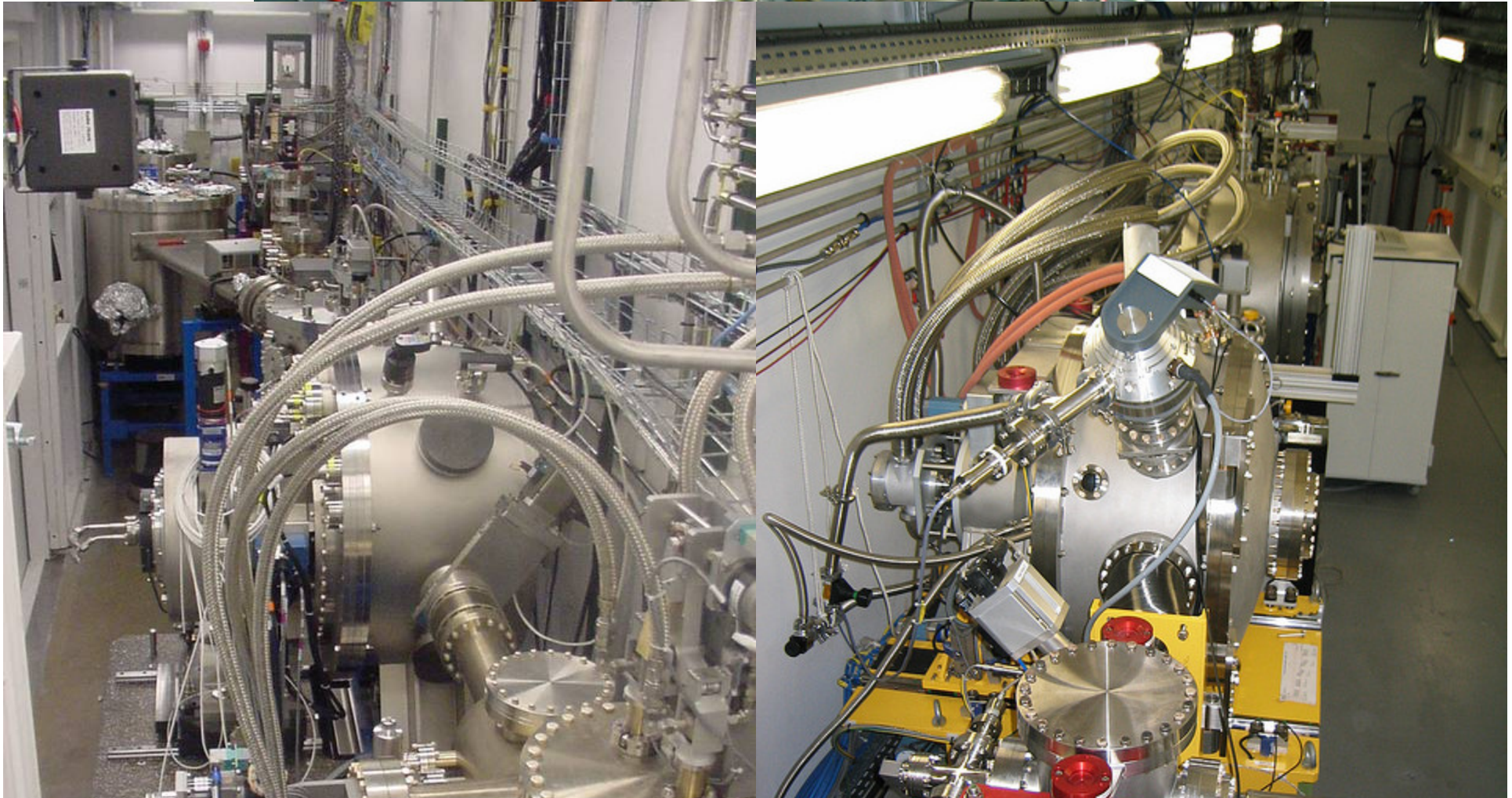
How to correct the beam height variation



Constant beam-height
double crystal monochromator

+ Monochromator -outlook

BL13MPW@PF, KEK

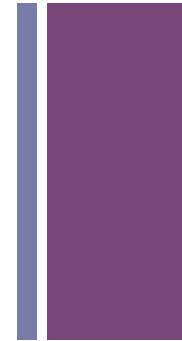


+ Beamline optics -strategy

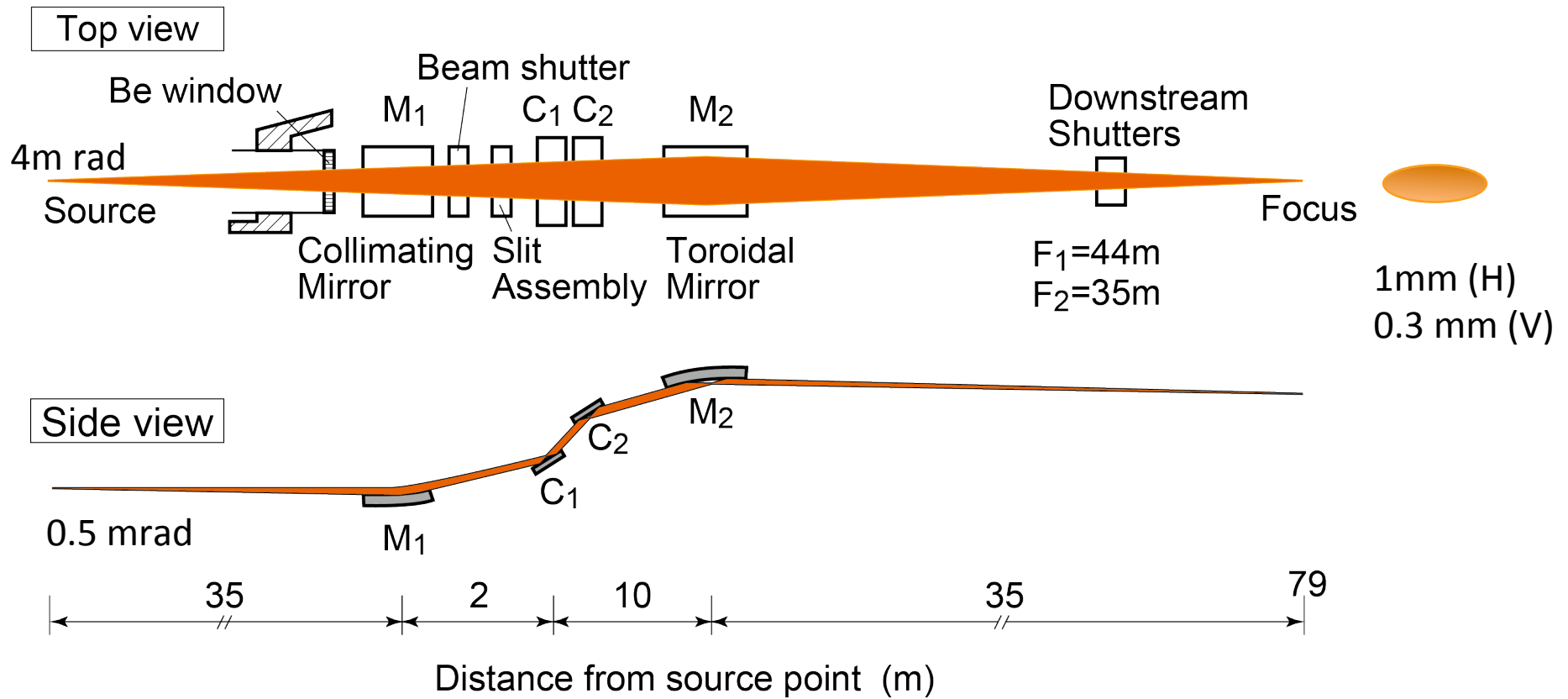
Wiggler@2G

Strategy: 1:1 mirror focusing or 3:1 sagittal focusing

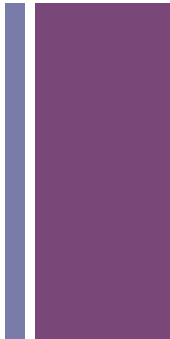
Emittance ≈ 30 nmrad



2 Crystal Monochromator

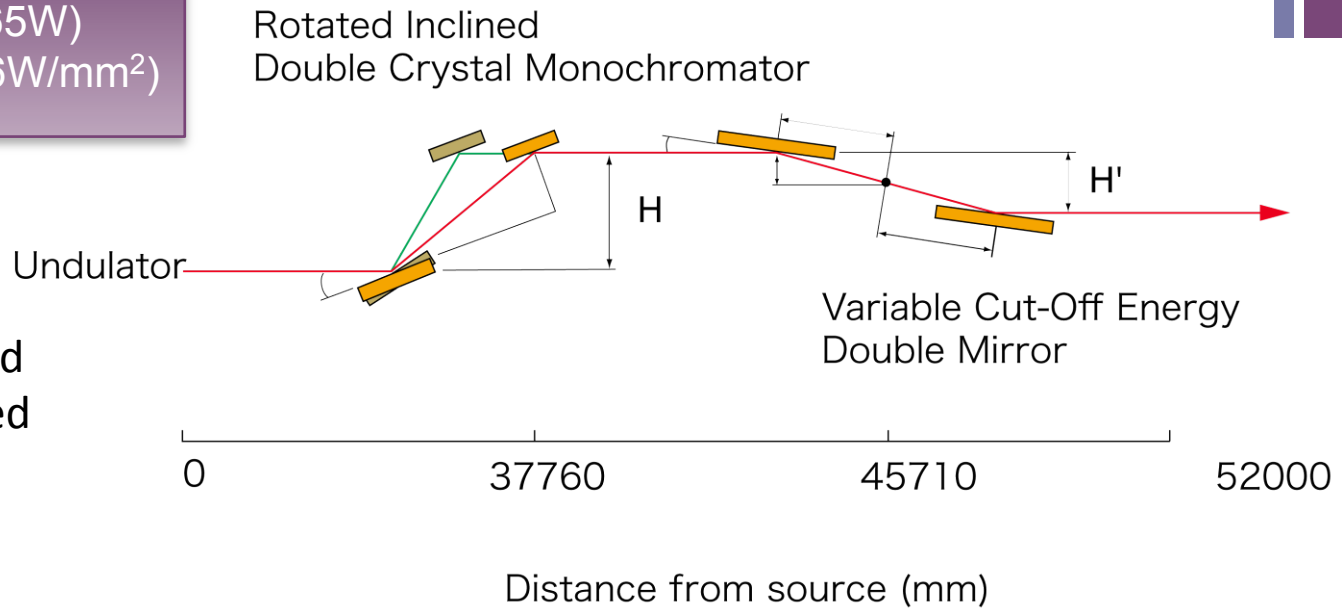


+ Optics of BL10XU (U032V) beamline@Spring-8



5-30 keV energy range
 1st & 3rd radiation (tunable)
 T.P. 12kW (465W)
 P.D. (76W/mm²)

Optics of SPring-8 BL10XU



Emittance: 3 nmrads
 Strategy: unfocused

Ray tracing results

